

Emission Spectrum of Hot HDO in the 380–2190 cm⁻¹ Region

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Received May 29, 2001; published online September 27, 2001

Fourier transform emission spectra were recorded using a mixture of H₂O and D₂O at a temperature of 1500°C. The spectra were recorded in three overlapping sections and cover the wavenumber range 380–2190 cm⁻¹. A total of 22 106 lines were measured, of which 60% are thought to belong to HDO. A total of 6430 HDO transitions are assigned, including the first transitions to the (040) vibrational state, with a term value of 5420.042 cm⁻¹. A total of 1536 new energy levels of HDO belonging to the (000), (010), (020), (030), and (040) states are presented, significantly extending the degree of rotational excitation compared to previous studies. © 2001 Elsevier Science

Key Words: water vapor; infrared spectrum; emission; line assignments; hot bands.

1. INTRODUCTION

In recent years there has been great progress in the experimental and theoretical understanding of the energy level structure of water (*1*). On the experimental side, new overtone spectra of water have been recorded in the visible and near UV regions (2–3). In the infrared, new spectra of hot water emission have now been measured (4–7) from 400 to 6000 cm⁻¹ in the laboratory. Hot water vapor lines have also been seen in absorption in the spectra of sunspots (4, 6, 8) (“Water on the Sun”). On the theoretical side, the availability of high quality *ab initio* potential energy surfaces (9) and the direct variational calculation of vibration–rotation energy levels have revolutionized the analysis of water spectra (10).

Progress for the HDO isotopomer has been much less satisfactory. In particular, no spectra of hot HDO have been analyzed to date. We report such observations in this paper.

The HDO molecule is of interest for a number of reasons. Because H₂O is so abundant in our atmosphere, HDO can be detected readily in atmospheric absorption spectra using the sun as a source (11). Astrophysicists are also interested in HDO because nearly all of the deuterium now in the Universe was formed in the Big Bang (12). The D/H ratio is thus an important parameter with cosmological significance. The ratio D/H can be determined for objects such as comets from the relative HDO to H₂O abundances (13).

In molecular physics, the HDO energy levels can be used to study the breakdown of the Born–Oppenheimer approximation. At the moment, this breakdown is the largest source of error in the calculation of water vibration–rotation energy levels from an *ab initio* potential energy surface. The potential surface can be corrected empirically (9–10), but recently Schwenke (14) calculated *ab initio* a complete set of corrections for the breakdown of the Born–Oppenheimer approximation in water. The experimental energy levels of both HDO and H₂O are needed for comparison with theory. The HDO molecule is also a popular molecule for mode-selective laser chemistry by dissociation (e.g., 15).

The infrared spectra of HDO were first measured by Benedict *et al.* (16) in 1956, followed by work in France (e.g., 17). Since then there has been considerable additional work, mainly by Toth and co-workers (18–23) on the room temperature infrared absorption spectra. Toth has recently published a list of the energy levels for the (000), (010), (100), and (020) vibrational states of HDO. Note that we adopt the traditional labeling convention of ν_1 for the OD stretching mode and for the ν_3 OH stretching mode, in spite of the Mulliken convention (24). Other infrared measurements include the detection of several transitions near 1 μm by Bykov *et al.* (25, 26) and the 300–000 and 111–000 bands by Hu *et al.* (27).

The pure rotational lines of HDO were measured in the sub-millimeter wave region by Messer *et al.* (28) and by Baskakov *et al.* (29). To higher frequency, the pure rotational transitions in the regions 20–350 cm⁻¹ (30) and 110–500 cm⁻¹ (31) were recorded by Fourier transform absorption spectroscopy.

The most highly excited levels of HDO have been recorded by overtone spectroscopy in the visible and near-UV regions.

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Most notably, Campargue and co-workers (32–36) have recently used the ultrasensitive technique of intracavity laser absorption spectroscopy to measure a number of overtone bands. Our contribution in this area has been to record long-path Fourier transform absorption spectra in the region 16 300–22 800 cm^{-1} (37). The most highly excited vibrational state known to date is (007) with a band origin at 22 625.50 cm^{-1} (37). Other experiments on HDO are the intracavity Fourier transform measurements of the 500–000 transition (38), as well as Russian intracavity measurements of the (003) and (005) levels (39).

2. EXPERIMENTAL DETAILS

The hot HDO emission spectra were recorded on November 6, 1997 at the University of Waterloo with a Bruker IFS 120HR Fourier transform spectrometer. The spectrometer was operated with a KBr beamsplitter and either a Si : B or a HgCdTe detector. The spectra reported here in the region 350–2200 cm^{-1} were recorded in three pieces. The section 350–750 cm^{-1} used a liquid-He-cooled Si : B detector and a cold longwave pass filter at 750 cm^{-1} . A separate cold bandpass filter was used to cover the region 750–1300 cm^{-1} . The region 1200–2200 cm^{-1} was recorded with a HgCdTe detector and an uncooled 2200 cm^{-1} longwave pass filter. Forty-five scans were co-added with the Si : B detector and forty with the HgCdTe detector at a resolution of 0.01 cm^{-1} .

A KRS-5 window was used on the emission port of the spectrometer. The water vapor was heated in the center of a 1-m-long, 5-cm-diameter alumina tube sealed with cooled KRS-5 windows. The tube was placed inside a furnace and heated to 1500°C. A slow flow of water vapor through the cell was maintained at a pressure of 2.5 Torr. An equimolar mixture of H₂O and D₂O liquids was used to provide the vapor. The thermal emission from the cell was focused into the emission port of the spectrometer with an off-axis parabolic mirror. The lines were measured with the PC-Decomp program of Brault and has an estimated accuracy of $\pm 0.001 \text{ cm}^{-1}$ for strong unblended lines. The spectrum, however, was very dense with H₂O, D₂O and HDO lines present.

The three spectra analyzed for this paper have lines in the regions 380–746, 750–1249, and 1250–2180 cm^{-1} . There were enough strong common lines in the two higher wavenumber spectra to put them on a common wavenumber scale and then calibrate them with the water lines reported in Polyansky *et al.* (40). Because of a lack of strong common lines between the two lower wavenumber regions, the lines in the region 380–746 cm^{-1} were just calibrated with our previous measurements (40) in this region. This means that in the region 750–2190 cm^{-1} our lines have a wavenumber scale that is in excellent agreement with that of Toth (41, 42), but the region 380–746 cm^{-1} is on a scale slightly different from that of Toth. Fortunately, this difference is less than 0.001 cm^{-1} , our estimated absolute accuracy.

3. LINE ASSIGNMENTS

A total of 22 106 lines were measured in the emission spectrum. Of course not all of these transitions correspond to HDO and before starting detailed analysis of these lines it was necessary to eliminate those due to H₂O and D₂O. The H₂O lines were identified by comparison with previously published (5, 6) hot emission spectra. For D₂O similar comparisons were made with a “pure” D₂O emission spectrum recorded in Waterloo which is yet to be fully analyzed; in practice this D₂O spectrum contained approximately 10% HDO. As the intensities differ in the three regions of the HDO spectrum, these regions were analysed separately. Lines were identified as H₂O or D₂O by matching both frequency and intensity. This is because the line density of the spectra is such that inevitably some HDO lines coincide with lines from the other isotopomers. We identified 4155 H₂O lines (322 also belonging to HDO) and 5423 D₂O (447 also HDO). H₂O and D₂O lines are marked in the full linelist which is given in the supplementary data for this article.

Energy levels for the ground, (010), (020), (030), (100), (001), and (110) vibrational states of HDO with low J and K_a values have been given by Toth (18, 19, 21, 22). These were used to conduct an initial analysis of the spectrum to identify “trivial” assignments were both upper and lower energy levels were already known.

To identify transitions involving previously unobserved energy levels in the system it was necessary to use variational predictions to aid the assignment process. There is no linelist available to us for hot HDO, but Partridge and Schwenke (9) have computed a room temperature linelist with states extending to energies about 8000 cm^{-1} above the ground state. This linelist was transformed to a temperature of 1800 K using Boltzmann statistics. The spectral region 380–2190 cm^{-1} covers both pure rotational transitions of HDO, particularly with high J or K_a , and the first bending fundamental, with a band origin at 1403.48 cm^{-1} , as well as associated hot bands.

Trivial assignments and other isotopomers having been eliminated from our list of transitions, the unassigned lines were analyzed using a computer program. Candidate transitions were identified using the variational linelist and then confirmed, or otherwise, by the presence, or not, of the appropriate combination difference transitions. In this way we were able to identify numerous new transitions involving known vibrational states and also to identify 190 transitions involving the (040) bending state.

The Partridge and Schwenke linelist (9) proved to be too restricted for the analysis of hot rotational levels. However, they also provided (43) energy levels for higher states of HDO. These were used to provide estimates of frequencies for pure rotational transitions which were then used to seed a further search.

TABLE 1
Summary of Assigned HDO
Emission Lines

Transition	Number
Pure rotations	
000-000	516
010-010	372
020-020	170
030-030	77
Bending modes	
010-000	1723
020-010	1676
030-020	1011
040-030	74
110-010	362
011-001	123
Difference bands	
100-010	65
110-020	132

We have assigned 1146 pure HDO rotational transitions spanning all HDO states up to (110). A total of 5085 transitions were assigned to bending transitions which involve a change of one quantum in the ν_2 mode. In addition 199 transitions were assigned to assorted difference bands. It has already been found that these difference bands are common in the emission spectrum of hot H₂O (5, 44). Table 1 gives a summary of all bands for which more than 10 transitions were assigned.

Experimental energy levels, which can be calculated by linking known energy levels and newly assigned transitions, are an important product of our analysis. This work has more than tripled the number of known energy levels for the states (000), (010), (020), (030) and (040). Tables 2 and 3 present energy levels for the first four of these states. The lowest J levels have been omitted from these tabulations as they have been well-determined previously (18, 19). Our results for the previously unobserved (040) state are given in Table 4.

Altogether we have assigned 6430 HDO transitions in the emission spectrum. This number represents approximately half the transitions which we identify as belonging to HDO. There is no doubt that many of the unassigned lines in the lower wavenumber portion of the spectrum belong to pure rotational transitions of vibrational states with excited stretching modes. Analysis of these states is best conducted in conjunction with analysis of associated vibrational transitions involving these stretching modes. Such transitions lie to somewhat higher frequency than the spectra reported here. An emission spectrum covering this region has been recorded in Waterloo and will be analyzed together with the unassigned transitions remaining from the present study.

TABLE 2
Experimental Term Values in cm⁻¹ for the Ground and
(010) Vibrational States of HDO

J	K_a	K_c	(000)		(010)	
11	0	11	916.029	a	2316.312	a
11	1	11	916.124	a	2316.446	a
11	1	10	1046.474	a	2459.361	a
11	2	10	1049.124	a	2462.887	a
11	2	9	1141.691	a	2561.580	a
11	3	9	1164.510	a	2589.750	a
11	3	8	1206.754	a	2632.660	a
11	4	8	1278.438	a	2716.170	a
11	4	7	1287.239	a	2724.463	a
11	5	7	1410.566	a	2863.369	a
11	5	6	1411.319	a	2864.036	a
11	6	6	1570.061	a	3040.100	a
11	6	5	1570.095	a	3040.129	a
11	7	5	1757.340	a	3246.058	a
11	7	4	1757.341	a	3246.059	a
11	8	4	1971.030	a	3479.363	a
11	8	3	1971.030	a	3479.363	a
11	9	3	2209.906	a	3738.413	a
11	9	2	2209.906	a	3738.413	a
11	10	2	2472.974	a	4021.937	d
11	10	1	2472.974	a	4021.937	d
11	11	1	2759.426	d	4328.919	d
11	11	0	2759.426	d	4328.919	d
12	0	12	1075.715	a	2474.835	a
12	1	12	1075.762	a	2474.904	a
12	1	11	1220.028	a	2633.352	a
12	2	11	1221.536	a	2635.417	a
12	2	10	1331.217	a	2752.926	a
12	3	10	1347.119	a	2773.107	a
12	3	9	1405.125	a	2832.490	a
12	4	9	1465.831	a	2904.291	a
12	4	8	1481.443	a	2919.136	a
12	5	8	1598.066	a	3051.433	a
12	5	7	1599.798	a	3052.969	a
12	6	7	1756.284	a	3226.829	a
12	6	6	1756.384	a	3226.912	a
12	7	6	1942.372	a	3431.571	a
12	7	5	1942.375	a	3431.573	a
12	8	5	2155.016	a	3663.801	a
12	8	4	2155.016	a	3663.801	a
12	9	4	2392.886	a	3921.804	a
12	9	3	2392.886	a	3921.804	a
12	10	3	2654.916	d	4204.229	d
12	10	2	2654.916	d	4204.229	d
12	11	2	2940.256	d	4510.028	d
12	11	1	2940.256	d	4510.028	d
12	12	1	3248.227	d	4838.388	d
12	12	0	3248.227	d	4838.388	d
13	0	13	1247.964	a	2645.775	a
13	1	13	1247.988	a	2645.811	a
13	1	12	1405.818	a	2819.396	a
13	2	12	1406.656	a	2820.577	a
13	2	11	1532.728	a	2956.268	a
13	3	11	1543.243	a	2970.013	a
13	3	10	1618.698	a	3047.832	a

^a From Toth (18).

^d Levels treated as degenerate.

TABLE 2—Continued

<i>J</i>	<i>K_a</i>	<i>K_c</i>	(000)		(010)	
13	4	10	1668.095	a	3107.381	a
13	4	9	1693.446	a	3131.745	a
13	5	9	1801.207	a	3255.187	a
13	5	8	1804.840	a	3258.418	a
13	6	8	1958.125	a	3429.192	a
13	6	7	1958.383	a	3429.410	a
13	7	7	2142.829	a	3632.524	a
13	7	6	2142.841	a	3632.533	a
13	8	6	2354.274	a	3863.529	a
13	8	5	2354.274	a	3863.529	a
13	9	5	2591.018	a	4120.363	d
13	9	4	2591.018	a	4120.363	d
13	10	4	2851.907	a	4401.589	d
13	10	3	2851.907	a	4401.589	d
13	11	3	3136.039	d	4706.101	d
13	11	2	3136.039	d	4706.101	d
13	12	2	3442.703	d	5033.060	d
13	12	1	3442.703	d	5033.060	d
13	13	1	3771.349	d	5381.836	d
13	13	0	3771.349	d	5381.836	d
14	0	14	1432.747	a	2829.106	a
14	1	14	1432.758	a	2829.125	a
14	1	13	1603.905	a	3017.575	a
14	2	13	1604.362	a	3018.240	a
14	2	12	1745.896	a	3171.148	a
14	3	12	1752.537	a	3180.106	a
14	3	11	1846.401	a	3277.601	a
14	4	11	1884.843	a	3325.047	a
14	4	10	1922.888	a	3362.052	a
14	5	10	2019.821	a	3474.478	a
14	5	9	2026.831	a	3480.745	a
14	6	9	2175.564	a	3647.170	a
14	6	8	2176.177	a	3647.687	a
14	7	8	2358.688	a	3848.900	a
14	7	7	2358.721	a	3848.926	a
14	8	7	2568.783	a	4078.518	d
14	8	6	2568.783	a	4078.518	d
14	9	6	2804.271	a	4334.055	d
14	9	5	2804.271	a	4334.055	d
14	10	5	3063.912	a	4613.975	d
14	10	4	3063.912	a	4613.975	d
14	11	4	3346.714	d	4917.099	d
14	11	3	3346.714	d	4917.099	d
14	12	3	3652.000	d	5242.558	d
14	12	2	3652.000	d	5242.558	d
14	13	2	3979.121	d	5589.694	d
14	13	1	3979.121	d	5589.694	d
14	14	1	4327.677	d	5958.044	d
14	14	0	4327.677	d	5958.044	d
15	0	15	1630.024	a	3024.796	a
15	1	15	1630.029	a	3024.805	a
15	1	14	1814.319	a	3227.938	a
15	2	14	1814.565	a	3228.307	a
15	2	13	1970.655	a	3397.438	a
15	3	13	1974.708	a	3403.071	a
15	3	12	2087.137	a	3520.662	a
15	4	12	2115.644	a	3556.859	a
15	4	11	2169.061	a	3609.422	a
15	5	11	2253.675	a	3709.077	a
15	5	10	2266.175	a	3720.348	a

TABLE 2—Continued

<i>J</i>	<i>K_a</i>	<i>K_c</i>	(000)		(010)	
15	6	10	2408.557	a	3880.715	a
15	6	9	2409.903	a	3881.850	a
15	7	9	2589.973	a	4080.660	
15	7	8	2590.057	a	4080.742	
15	8	8	2798.509	a	4308.719	
15	8	7	2798.513	a	4308.722	
15	9	7	3032.606	a	4562.833	d
15	9	6	3032.606	a	4562.833	d
15	10	6	3290.888	d	4841.335	d
15	10	5	3290.888	d	4841.335	d
15	11	5	3572.309	d	5142.978	d
15	11	4	3572.309	d	5142.978	d
15	12	4	3876.050	d	5466.838	d
15	12	3	3876.050	d	5466.838	d
15	13	3	4201.580	d	5812.270	d
15	13	2	4201.580	d	5812.270	d
15	14	2	4548.373	d	6178.683	d
15	14	1	4548.373	d	6178.683	d
15	15	1	4916.149	d	6565.872	d
15	15	0	4916.149	d	6565.872	d
16	0	16	1839.754	a	3232.804	a
16	1	16	1839.757	a	3232.809	a
16	1	15	2037.056	a	3450.501	a
16	2	15	2037.188	a	3450.705	a
16	2	14	2207.117	a	3635.201	a
16	3	14	2209.513		3638.648	a
16	3	13	2339.944		3776.003	
16	4	13	2360.075		3802.374	
16	4	12	2430.992	a	3872.934	
16	5	12	2502.437	a	3958.678	a
16	5	11	2523.127		3977.541	
16	6	11	2657.013	a	4129.736	
16	6	10	2659.764		4132.058	
16	7	10	2836.617		4327.808	
16	7	9	2836.823	a	4327.975	
16	8	9	3043.420	a	4554.094	
16	8	8	3043.431	a	4554.090	
16	9	8	3275.978	a	4806.638	d
16	9	7	3275.978	a	4806.638	d
16	10	7	3532.781	d	5083.618	d
16	10	6	3532.781	d	5083.618	d
16	11	6	3812.705	d	5383.681	d
16	11	5	3812.705	d	5383.681	d
16	12	5	4114.878	d	5705.846	d
16	12	4	4114.878	d	5705.846	d
16	13	4	4438.651	d	6049.390	d
16	13	3	4438.651	d	6049.390	d
16	14	3	4783.618	d	6413.860	d
16	14	2	4783.618	d	6413.860	d
16	15	2	5149.378	d	6798.868	d
16	15	1	5149.378	d	6798.868	d
16	16	1	5535.745	d	7204.241	d
16	16	0	5535.745	d	7204.241	d
17	0	17	2061.897	a	3453.092	a
17	1	17	2061.898	a	3453.094	a
17	1	16	2272.103	a	3685.259	a
17	2	16	2272.172	a	3685.371	a
17	2	15	2455.379	a	3884.565	a
17	3	15	2456.767		3886.630	a
17	3	14	2604.144		4042.765	

TABLE 2—Continued

<i>J</i>	<i>K_a</i>	<i>K_c</i>	(000)	(010))
17	4	14	2617.726	4061.184
17	4	13	2707.595	a 4151.511
17	5	13	2765.744	4222.914
17	5	12	2797.620	4252.371
17	6	12	2920.783	a 4394.100
17	6	11	2926.032	4398.558
17	7	11	3098.595	4590.252
17	7	10	3099.058	a 4590.627
17	8	10	3303.477	4814.593
17	8	9	3303.503	4814.593
17	9	9	3534.342	d 5065.396 d
17	9	8	3534.342	d 5065.396 d
17	10	8	3789.538	d 5340.750 d
17	10	7	3789.538	d 5340.750 d
17	11	7	4067.863	d 5639.141 d
17	11	6	4067.863	d 5639.141 d
17	12	6	4368.360	d 5959.524 d
17	12	5	4368.360	d 5959.524 d
17	13	5	4690.352	d 6301.143 d
17	13	4	4690.352	d 6301.143 d
17	14	4	5033.327	d 6663.518 d
17	14	3	5033.327	d 6663.518 d
17	15	3	5397.011	d 7046.250 d
17	15	2	5397.011	d 7046.250 d
17	16	2	5781.098	d 7449.132 d
17	16	1	5781.098	d 7449.132 d
17	17	1	6185.485	d 7872.136 d
17	17	0	6185.485	d 7872.136 d
18	0	18	2296.384	a 3685.619 a
18	1	18	2296.384	a 3685.619 a
18	1	17	2519.416	a 3932.188 a
18	2	17	2519.452	a 3932.250 a
18	2	16	2715.533	4145.632
18	3	16	2716.330	4146.842
18	3	15	2879.423	4320.513
18	4	15	2888.238	4332.875
18	4	14	2997.706	4443.997
18	5	14	3043.157	4501.367
18	5	13	3089.176	4544.524
18	6	13	3199.653	4673.591
18	6	12	3209.060	4681.624
18	7	12	3375.825	4867.944
18	7	11	3376.824	4868.733
18	8	11	3578.631	5090.161
18	8	10	3578.700	5090.218
18	9	10	3807.626	5339.075
18	9	9	3807.623	5339.124
18	10	9	4061.120	d 5612.657
18	10	8	4061.120	d 5612.668
18	11	8	4337.716	5909.282 d
18	11	7	4337.732	5909.282 d
18	12	7	4636.448	d 6227.795 d
18	12	6	4636.448	d 6227.795 d
18	13	6	4956.561	d 6567.408 d
18	13	5	4956.561	d 6567.408 d
18	14	5	5297.521	d 6927.609 d
18	14	4	5297.521	d 6927.609 d
18	15	4	5658.963	d 7307.960 d
18	15	3	5658.963	d 7307.960 d
18	16	3	6040.706	d 7708.295 d

TABLE 2—Continued

<i>J</i>	<i>K_a</i>	<i>K_c</i>	(000)	(010))
18	16	2	6040.706	d 7708.295 d
18	17	2	6442.540	d 8128.485 d
18	17	1	6442.540	d 8128.485 d
18	18	1	6864.423	d 8568.592 d
18	18	0	6864.423	d 8568.592 d
19	0	19	2543.209	a 3930.315 a
19	1	19	2543.209	a 3930.315 a
19	1	18	2778.949	a 4191.256 a
19	2	18	2778.969	a 4191.291 a
19	2	17	2987.607	4418.476
19	3	17	2988.058	4419.190
19	3	16	3165.744	4609.095
19	4	16	3171.279	4617.121
19	4	15	3300.211	4749.217
19	5	15	3334.237	4793.553
19	5	14	3396.998	4853.290
19	6	14	3493.330	4967.941
19	6	13	3509.153	4981.591
19	7	13	3668.238	5160.756
19	7	12	3670.218	5162.347
19	8	12	3868.826	5380.715
19	8	11	3868.969	5380.851
19	9	11	4095.786	5627.594
19	9	10	4095.789	5627.622
19	10	10	4347.409	d 5899.252
19	10	9	4347.409	d 5899.260
19	11	9	4622.221	6194.051
19	11	8	4622.256	6194.016
19	12	8	4919.073	d 6510.577 d
19	12	7	4919.073	d 6510.577 d
19	13	7	5237.223	d 6848.113 d
19	13	6	5237.223	d 6848.113 d
19	14	6	5576.072	d 7206.073 d
19	14	5	5576.072	d 7206.073 d
19	15	5	5935.256	d 7583.951 d
19	15	4	5935.256	d 7583.951 d
19	16	4	6314.483	d 7981.602 d
19	16	3	6314.483	d 7981.602 d
19	17	3	6713.698	d 8398.940 d
19	17	2	6713.698	d 8398.940 d
19	18	2	7132.751	d 8835.273 d
19	18	1	7132.751	d 8835.273 d
19	19	1	7571.635	d 9292.624 d
19	19	0	7571.635	d 9292.624 d
20	0	20	2802.260	d 4187.166 d
20	1	20	2802.260	d 4187.166 d
20	1	19	3050.657	d 4462.397
20	2	19	3050.657	d 4462.420
20	2	18	3271.646	4703.124
20	3	18	3271.881	4703.542
20	3	17	3463.200	4908.538
20	4	17	3466.595	4913.621
20	4	16	3614.113	5066.114
20	5	16	3638.520	5099.072
20	5	15	3720.010	5177.710
20	6	15	3801.480	5276.814
20	6	14	3826.480	5298.647
20	7	14	3975.666	5468.553
20	7	13	3979.411	5471.580
20	8	13	4173.984	5686.167

TABLE 2—Continued

J	K_a	K_c	(000)	(010)
20	8	12	4174.310	5686.451
20	9	12	4398.724	5930.817
20	9	11	4398.751	5930.876
20	10	11	4648.370	d 6200.454
20	10	10	4648.370	d 6200.486
20	11	10	4921.249	6492.761
20	11	9	4921.298	6493.259
20	12	9	5216.184	6807.780
20	12	8	5216.183	6807.780
20	13	8	5532.262	d 7143.146
20	13	7	5532.262	d 7143.146
20	14	7	5868.922	d 7498.846
20	14	6	5868.922	d 7498.846
20	15	6	6225.765	d 7874.153
20	15	5	6225.765	d 7874.153
20	16	5	6602.449	d 8269.020
20	16	4	6602.449	d 8269.020
20	17	4	6998.872	d 8683.378
20	17	3	6998.872	d 8683.378
20	18	3	7415.020	d 9117.080
20	18	2	7415.020	d 9117.080
20	19	2	7850.792	d
20	19	1	7850.792	d
20	20	1	8306.233	d
20	20	0	8306.233	d
21	0	21	3073.522	d 4456.096
21	1	21	3073.522	d 4456.096
21	1	20	3334.489	d 4745.621
21	2	20	3334.489	d 4745.655
21	2	19	3567.588	4999.574
21	3	19	3567.717	4999.845
21	3	18	3771.939	5219.015
21	4	18	3773.988	5222.179
21	4	17	3938.799	5393.863
21	5	17	3955.624	5417.441
21	5	16	4057.068	5516.611
21	6	16	4123.665	5599.847
21	6	15	4160.851	5632.839
21	7	15	4297.931	5791.130
21	7	14	4304.654	5796.611
21	8	14	4494.019	6006.537
21	8	13	4494.722	6007.002
21	9	13	4716.394	6248.737
21	9	12	4716.430	6248.812
21	10	12	4963.892	6516.115
21	10	11	4963.865	6516.170
21	11	11	5234.753	6806.751
21	11	10	5234.811	6806.916
21	12	10	5527.668	7119.314
21	12	9	5527.638	7119.294
21	13	9	5841.622	d 7452.436
21	13	8	5841.622	d 7452.436
21	14	8	6175.990	d 7805.845
21	14	7	6175.990	d 7805.845
21	15	7	6530.456	d 8178.497
21	15	6	6530.456	d 8178.497
21	16	6	6904.470	d 8570.474
21	16	5	6904.470	d 8570.474
21	17	5	7298.083	d 8981.761
21	17	4	7298.083	d 8981.761

TABLE 2—Continued

J	K_a	K_c	(000)	(010)
21	18	4	7711.152	d 9412.196
21	18	3	7711.152	d 9412.196
21	19	3	8143.730	d 9862.164
21	19	2	8143.730	d 9862.164
21	20	2	8595.807	d
21	20	1	8595.807	d
22	0	22	3356.925	d 4737.062
22	1	22	3356.925	d 4737.062
22	1	21	3630.374	d 5040.850
22	2	21	3630.374	d 5040.747
22	2	20	3875.405	5307.810
22	3	20	3875.470	5308.036
22	3	19	4092.042	5540.636
22	4	19	4093.277	5542.561
22	4	18	4273.922	5731.948
22	5	18	4285.128	5748.263
22	5	17	4406.977	5868.944
22	6	17	4459.463	5936.508
22	6	16	4511.727	5983.725
22	7	16	4634.736	6128.461
22	7	15	4646.194	6137.661
22	8	15	4828.835	6341.365
22	8	14	4830.210	6342.450
22	9	14	5048.678	6581.137
22	9	13	5048.896	6581.252
22	10	13	5293.884	6845.751
22	10	12	5293.801	6845.827
22	11	12	5562.592	7134.677
22	11	11	5562.690	7134.862
22	12	11	5853.317	7445.060
22	12	10	5853.387	7445.120
22	13	10	6165.160	d 7775.878
22	13	9	6165.161	d 7775.878
22	14	9	6497.213	d 8127.124
22	14	8	6497.213	d 8127.271
22	15	8	6849.654	d 8496.873
22	15	7	6849.654	d 8496.873
22	16	7	7220.487	d 8885.837
22	16	6	7220.487	d 8885.837
22	17	6	7611.189	d 9294.019
22	17	5	7611.189	d 9294.019
22	18	5	8021.164	d 9721.085
22	18	4	8021.164	d 9721.085
22	19	4	8450.371	d
22	19	3	8450.371	d
22	20	3		10632.566
22	20	2		10632.566
22	21	2	9366.836	d
22	21	1	9366.836	d
23	0	23	3652.398	d 5030.012
23	1	23	3652.398	d 5030.012
23	1	22	3938.265	d 5347.989
23	2	22	3938.265	d 5348.017
23	2	21	4195.052	5627.771
23	3	21	4195.107	5627.642
23	3	20	4423.568	5873.463
23	4	20	4424.292	5874.594
23	4	19	4619.446	6080.213
23	5	19	4626.722	6091.177
23	5	18	4768.576	6233.324

TABLE 2—Continued

<i>J</i>	<i>K_a</i>	<i>K_c</i>	(000)	(010))
23	6	18	4808.411	6286.414
23	6	17	4878.296	6350.622
23	7	17	4985.818	6479.813
23	7	16	5004.291	6494.921
23	8	16	5178.253	6690.747
23	8	15	5180.822	6692.876
23	9	15	5395.498	6927.921
23	9	14	5395.765	6928.185
23	10	14	5638.233	7190.978
23	10	13	5638.455	7191.122
23	11	13	5904.704	7476.729
23	11	12	5904.816	7476.997
23	12	12	6193.249	7784.920
23	12	11	6193.282	7784.816
23	13	11	6502.823	8113.377
23	13	10	6502.860	8113.337
23	14	10	6832.444	d 8463.318
23	14	9	6832.444	d 8463.421
23	15	9	7181.439	d 8829.215
23	15	8	7181.439	d 8829.215
23	16	8	7550.428	d 9216.976
23	16	7	7550.428	d 9216.976
23	17	7	7938.136	d 9620.066
23	17	6	7938.136	d 9620.066
23	18	6	8344.912	d
23	18	5	8344.912	d
23	20	4	9213.862	d
23	20	3	9213.862	d
23	21	3		11429.137
23	21	2		11429.137
24	0	24	3959.891	5334.785
24	1	24	3959.891	5334.863
24	1	23	4258.076	5667.067
24	2	23	4258.081	5667.101
24	2	22	4526.455	5959.349
24	3	22	4526.485	5959.423
24	3	21	4766.537	6217.584
24	4	21	4766.968	6218.237
24	4	20	4975.506	6438.692
24	5	20	4980.108	6445.890
24	5	19	5140.984	6608.881
24	6	19	5170.049	6649.089
24	6	18	5259.364	6732.807
24	7	18	5350.860	6845.050
24	7	17	5379.016	6868.741
24	8	17	5542.104	7053.298
24	8	16	5546.741	7058.249
24	9	16	5756.701	7288.980
24	9	15	5757.217	7289.506
24	10	15	5996.865	7549.341
24	10	14	5996.999	7549.555
24	11	14	6260.933	7832.690
24	11	13	6261.034	7833.146
24	12	13	6547.199	8138.808
24	12	12	6547.251	8138.644
24	13	12	6854.498	8464.738
24	13	11	6854.372	8464.716
24	14	11	7181.611	8806.790
24	14	10	7181.617	8807.590
24	15	10	7528.079	d 9175.441

TABLE 2—Continued

<i>J</i>	<i>K_a</i>	<i>K_c</i>	(000)	(010))
24	15	9	7528.079	d 9175.441
24	16	9	7894.226	d
24	16	8	7894.226	d 9558.701
24	17	8	8278.846	d
24	17	7	8278.846	d
24	18	7	8682.347	d
24	18	6	8682.347	d
24	19	6	9104.665	d
24	19	5	9104.665	d
25	0	25	4279.321	5651.141
25	1	25	4279.320	5651.555
25	1	24	4589.759	5997.975
25	2	24	4589.762	5998.035
25	2	23	4869.569	d 6302.605
25	3	23	4869.569	d 6302.679
25	3	22	5120.929	6572.948
25	4	22	5121.170	6573.279
25	4	21	5342.191	6807.427
25	5	21	5345.054	6812.044
25	5	20	5523.459	6994.663
25	6	20	5543.933	7024.132
25	6	19		7128.614
25	7	19	5729.143	7224.271
25	7	18	5770.084	7258.859
25	8	18	5920.136	7433.214
25	8	17	5928.097	7438.777
25	9	17	6132.213	
25	9	16	6133.163	7665.121
25	10	16	6369.613	7921.830
25	10	15	6369.775	
25	11	15	6631.188	8202.490
25	11	14	6631.319	8203.137
25	12	14	6915.017	8506.605
25	12	13	6915.132	8506.267
25	13	13		8829.914
25	13	12		8829.871
25	14	12	7544.584	d 9170.915
25	14	11	7544.584	d 9171.232
25	15	11		9535.491
25	15	10		9535.481
25	16	10	8251.751	9915.396
25	16	9	8251.683	9915.396
25	17	9	8633.244	d
25	17	8	8633.244	d
26	0	26	4610.681	5980.020
26	1	26	4610.695	5980.020
26	1	25	4933.236	6340.678
26	2	25	4933.225	6340.881
26	2	24	5224.284	6657.395
26	3	24	5224.285	6657.482
26	3	23	5486.675	6939.529
26	4	23	5486.816	6939.677
26	4	22	5719.617	7186.554
26	5	22	5721.398	7189.433
26	5	21	5915.732	
26	6	21	5929.711	7410.985
26	6	20	6060.513	7537.166
26	7	20	6120.538	
26	7	19	6177.083	
26	8	19	6311.976	7824.851

TABLE 2—Continued

<i>J</i>	<i>K_a</i>	<i>K_c</i>	(000)	(010)
26	8	18	6325.097	7834.413
26	9	18	6521.824	8052.206
26	9	17	6523.580	
26	10	17	6756.370	
26	10	16	6756.649	
26	11	16	7015.293	
26	11	15	7015.471	
26	12	15	7296.663	
26	12	14	7296.823	
26	13	14	7599.052	
26	13	13	7599.055	
27	0	27	4953.766	d 6320.332 d
27	1	27	4953.766	d 6320.332 d
27	1	26	5288.457	6694.723
27	2	26	5288.452	6695.133
27	2	25	5590.581	d 7023.628
27	3	25	5590.566	d 7023.788
27	3	24	5863.717	7317.341
27	4	24	5863.808	7317.265
27	4	23	6107.851	7575.770
27	5	23	6108.923	7577.429
27	5	22	6317.700	
27	6	22	6327.007	
27	6	21	6478.228	
27	7	21	6524.436	
27	7	20	6598.915	
27	8	19	6737.922	
27	9	19	6925.321	
27	9	18	6928.513	
27	10	18	7156.989	
27	10	17	7157.462	
27	11	17	7413.162	
27	11	16	7413.377	
27	12	16	7691.922	
27	12	15	7692.139	
27	13	15	7992.035	
27	13	14	7991.856	
27	14	14	8311.383	d
27	14	13	8311.380	d
28	0	28	5308.590	d 6672.286 d
28	1	28	5308.590	d 6672.286 d
28	1	27	5655.240	7061.489
28	2	27	5655.231	7061.055
28	2	26	5968.269	d 7401.298
28	3	26	5968.269	d 7401.602
28	3	25	6251.986	7706.042
28	4	25	6252.026	7706.222
28	4	24	6506.866	7976.087
28	5	24	6507.530	7976.999
28	5	23	6729.440	
28	6	23	6735.531	
28	6	22	6906.086	
28	7	22	6940.441	
28	8	21	7136.548	
29	0	29	5675.066	d 7035.849 d
29	1	29	5675.066	d 7035.849 d
29	1	28		7438.591
29	2	28		7438.726
29	2	27	6357.327	d 7790.198
29	3	27	6357.327	d 7791.556

TABLE 2—Continued

<i>J</i>	<i>K_a</i>	<i>K_c</i>	(000)	(010)
29	3	26	6651.376	
29	4	26	6651.410	
29	4	25	6916.640	
29	5	25	6917.028	
29	5	24	7151.108	
29	6	24	7155.008	
30	0	30		7411.070 d
30	1	30		7411.070 d
30	2	28	6757.643	d
30	3	28	6757.643	d
30	3	27	7061.823	
30	4	27	7061.832	
30	4	26	7337.012	
30	5	26	7337.253	

TABLE 3

Experimental Term Values in cm^{-1} for the (020) and (030) Vibrational States of HDO

<i>J</i>	<i>K_a</i>	<i>K_c</i>	(020)	(030)
6	0	6	3089.126	a 4451.759 b
6	1	6	3092.811	a 4456.083 b
6	1	5	3154.965	a 4519.117 b
6	2	5	3186.626	a 4555.954 b
6	2	4	3206.336	a 4575.059 b
6	3	4	3291.072	a 4672.403 b
6	3	3	3292.976	a 4674.109 b
6	4	3	3425.685	a 4825.009 b
6	4	2	3425.748	a 4825.057 b
6	5	2	3594.668	a 5015.533 b
6	5	1	3594.669	a 5015.534 b
6	6	1	3795.879	a 5240.158 d
6	6	0	3795.879	a 5240.158 d
7	0	7	3185.579	a 4548.066 b
7	1	7	3187.760	a 4550.724 b
7	1	6	3268.346	a 4632.882 b
7	2	6	3292.830	a 4662.303 b
7	2	5	3324.490	a 4693.313 b
7	3	5	3401.218	a 4782.865 b
7	3	4	3405.847	a 4786.940 b
7	4	4	3535.818	a 4935.425 b
7	4	3	3536.047	a 4935.606 b
7	5	3	3704.195	a 5125.396 b
7	5	2	3704.199	a 5125.400 b
7	6	2	3904.897	a 5349.536 d
7	6	1	3904.898	a 5349.536 d
7	7	1	4135.476	a 5604.348 d
7	7	0	4135.476	a 5604.348 d
8	0	8	3294.197	a 4656.481 b
8	1	8	3295.443	a 4658.062 b
8	1	7	3395.168	a 4760.389 b

^a From Toth (18).

^b From Toth (21).

^c Level reassigned from Toth (21).

^d Levels treated as degenerate.

TABLE 3—Continued

<i>J</i>	<i>K_a</i>	<i>K_c</i>	(020)		(030)	
8	2	7	3413.012	a	4782.755	b
8	2	6	3459.535	a	4828.795	b
8	3	6	3526.894	a	4908.896	b
8	3	5	3536.354	a	4917.294	b
8	4	5	3661.860	a	5061.768	
8	4	4	3662.525	a	5062.297	
8	5	4	3829.486	a	5251.033	b
8	5	3	3829.508	a	5251.051	b
8	6	3	4029.533	a	5474.554	d
8	6	2	4029.533	a	5474.554	d
8	7	2	4259.517	a	5728.796	d
8	7	1	4259.517	a	5728.796	d
8	8	1	4517.274	a	6010.901	d
8	8	0	4517.274	a	6010.901	d
9	0	9	3415.036	a	4777.040	b
9	1	9	3415.732	a	4777.960	b
9	1	8	3534.570	a	4900.782	b
9	2	8	3546.853	a	4916.961	b
9	2	7	3610.681	a	4980.608	b
9	3	7	3667.776	a	5050.239	b
9	3	6	3684.868	a	5065.595	
9	4	6	3803.790	a	5204.030	b
9	4	5	3805.449	a	5205.350	b
9	5	5	3970.574	a	5392.467	c
9	5	4	3970.645	a	5392.524	c
9	6	4	4169.791	a	5615.242	
9	6	3	4169.790	a	5615.242	
9	7	3	4399.060	a	5868.765	d
9	7	2	4399.060	a	5868.765	d
9	8	2	4656.109	d	6150.151	d
9	8	1	4656.109	d	6150.151	d
9	9	1	4939.118	d	6457.181	d
9	9	0	4939.118	d	6457.181	d
10	0	10	3548.148	a	4909.788	b
10	1	10	3548.530	a	4910.315	b
10	1	9	3685.981	a	5053.412	b
10	2	9	3694.020	a	5064.591	b
10	2	8	3777.003	a	5148.026	b
10	3	8	3823.522	a	5206.573	
10	3	7	3851.399	a	5231.952	
10	4	7	3961.528	a	5362.130	
10	4	6	3965.184		5365.038	
10	5	6	4127.477	a	5549.718	
10	5	5	4127.688	a	5549.876	
10	6	5	4325.684	a	5771.521	d
10	6	4	4325.686	a	5771.521	d
10	7	4	4554.094	d	6024.234	d
10	7	3	4554.094	d	6024.234	d
10	8	3	4810.320	d	6304.800	d
10	8	2	4810.320	d	6304.800	d
10	9	2	5092.462	d	6610.936	d
10	9	1	5092.462	d	6610.936	d
10	10	1	5399.039	d	6940.950	d
10	10	0	5399.039	d	6940.951	d
11	0	11	3693.560	a	5054.751	b
11	1	11	3693.768	a	5055.050	b
11	1	10	3849.171	a	5217.919	b
11	2	10	3854.222	a	5225.344	b
11	2	9	3957.464	a	5330.063	
11	3	9	3993.756	a	5377.525	

TABLE 3—Continued

<i>J</i>	<i>K_a</i>	<i>K_c</i>	(020)		(030)	
11	3	8	4035.537	a	5416.346	
11	4	8	4134.914	a	5535.919	
11	4	7	4142.173	a	5541.765	
11	5	7	4300.205		5722.771	
11	5	6	4300.743		5723.160	
11	6	6	4497.212		5943.435	
11	6	5	4497.234		5943.453	
11	7	5	4724.597	d	6195.163	
11	7	4	4724.597	d	6195.164	
11	8	4	4979.884	d	6474.798	
11	8	3	4979.884	d	6474.798	
11	9	3	5261.060	d	6779.960	d
11	9	2	5261.060	d	6779.960	d
11	10	2	5566.586	d	7108.852	d
11	10	1	5566.586	d	7108.852	d
11	11	1	5895.293	d	7460.204	d
11	11	0	5895.293	d	7460.204	d
12	0	12	3851.280	a	5211.937	b
12	1	12	3851.391	a	5212.106	b
12	1	11	4024.142	a	5394.190	
12	2	11	4027.219	a	5398.980	
12	2	10	4151.015	a	5525.707	
12	3	10	4177.912	a	5562.713	
12	3	9	4236.551	a	5617.859	
12	4	9	4323.706		5725.215	
12	4	8	4336.861		5735.733	
12	5	8	4488.727		5911.820	
12	5	7	4489.966		5912.591	
12	6	7	4684.378		6130.957	
12	6	6	4684.447		6130.998	
12	7	6	4910.560		6381.545	
12	7	5	4910.562		6381.547	
12	8	5	5164.766	d	6660.119	
12	8	4	5164.766	d	6660.124	
12	9	4	5444.871	d	6964.188	
12	9	3	5444.871	d	6964.189	
12	10	3	5749.250	d	7291.935	d
12	10	2	5749.250	d	7291.935	d
12	11	2	6076.692	d	7641.921	d
12	11	1	6076.692	d	7641.921	d
12	12	1	6426.289	d	8013.230	d
12	12	0	6426.289	d	8013.230	d
13	0	13	4021.296	a	5381.322	b
13	1	13	4021.354	a	5381.423	b
13	1	12	4210.990	a	5582.254	
13	2	12	4212.824	a	5585.279	
13	2	11	4356.753	a	5734.016	
13	3	11	4375.881	a	5761.719	
13	3	10	4453.505		5835.820	
13	4	10	4527.591		5929.705	
13	4	9	4549.511		5948.035	
13	5	9	4692.963		6116.181	
13	5	8	4695.583		6118.087	
13	6	8	4887.169		6334.060	
13	6	7	4887.336		6334.173	
13	7	7	5111.967		6583.309	
13	7	6	5111.965		6583.317	
13	8	6	5364.940		6860.666	
13	8	5	5364.962		6860.702	
13	9	5	5643.861	d	7163.574	d

TABLE 3—Continued

<i>J</i>	<i>K_a</i>	<i>K_c</i>	(020)		(030)	
13	9	4	5643.861	d	7163.574	d
13	10	4	5946.981	d	7490.065	d
13	10	3	5946.981	d	7490.065	d
13	11	3	6273.070	d	7838.575	d
13	11	2	6273.070	d	7838.575	d
13	12	2	6621.146	d	8208.206	d
13	12	1	6621.146	d	8208.206	d
13	13	1	6990.499	d	8599.229	d
13	13	0	6990.499	d	8599.229	d
14	0	14	4203.589	a	5562.927	
14	1	14	4203.613	a	5562.982	
14	1	13	4409.821	a	5782.226	
14	2	13	4410.895	a	5784.103	
14	2	12	4574.064		5954.206	
14	3	12	4587.080		5974.167	
14	3	11	4685.335		6069.247	
14	4	11	4746.182		6149.058	
14	4	10	4780.039		6177.898	
14	5	10	4912.775		6336.283	
14	5	9	4917.900		6340.039	
14	6	9	5105.553		6552.689	
14	6	8	5105.951		6552.949	
14	7	8	5328.760		6800.405	
14	7	7	5328.778		6800.438	
14	8	7	5580.359		7076.506	
14	8	6	5580.361		7076.482	
14	9	6	5857.987	d	7378.061	d
14	9	5	5857.987	d	7378.061	d
14	10	5	6159.757		7703.206	d
14	10	4	6159.741		7703.206	d
14	11	4	6484.356	d	8050.141	d
14	11	3	6484.356	d	8050.141	d
14	12	3	6830.839	d	8417.949	d
14	12	2	6830.839	d	8417.949	d
14	13	2	7198.305	d	8807.787	d
14	13	1	7198.305	d	8807.787	d
14	14	1	7587.022	d	9215.047	d
14	14	0	7587.022	d	9215.047	d
15	0	15	4398.133	a	5756.723	
15	1	15	4398.144	a	5756.685	
15	1	14	4620.707	a	5994.169	
15	2	14	4621.328	a	5995.322	
15	2	13	4802.663		6185.804	
15	3	13	4811.219		6199.696	
15	3	12	4930.919		6317.066	
15	4	12	4979.068		6382.887	
15	4	11	5027.996		6425.431	
15	5	11	5147.962		6571.788	
15	5	10	5157.261		6578.695	
15	6	10	5339.468		6786.789	
15	6	9	5340.341		6787.367	
15	7	9	5560.912		7032.879	
15	7	8	5560.968		7032.852	
15	8	8	5810.921		7307.404	
15	8	7	5810.935		7307.400	
15	9	7	6087.162		7607.569	
15	9	6	6087.161		7607.554	
15	10	6	6387.476		7931.301	d
15	10	5	6387.476		7931.301	d
15	11	5	6710.528	d	8276.553	d
15	11	4	6710.528	d	8276.553	d

TABLE 3—Continued

<i>J</i>	<i>K_a</i>	<i>K_c</i>	(020)		(030)	
15	12	4	7055.295	d	8642.442	d
15	12	3	7055.295	d	8642.442	d
15	13	3	7420.797	d	9030.968	d
15	13	2	7420.797	d	9030.968	d
15	14	2	7807.703	d	9435.691	d
15	14	1	7807.703	d	9435.691	d
15	15	1	8213.975	d	9860.164	d
15	15	0	8213.975	d	9860.164	d
16	0	16	4604.889	a	5962.595	
16	1	16	4604.895	a	5962.616	
16	1	15	4843.687		6218.156	
16	2	15	4844.048		6218.862	
16	2	14	5042.533		6428.599	
16	3	14	5047.993		6437.996	
16	3	13	5189.148		6578.216	
16	4	13	5225.805		6630.897	
16	4	12	5292.562		6690.038	
16	5	12	5398.267		6822.431	
16	5	11	5414.029		6834.354	
16	6	11	5588.862		7036.262	
16	6	10	5590.624		7037.456	
16	7	10	5808.391			
16	7	9	5808.515		7280.540	
16	8	9	6056.728			
16	8	8	6056.729			
16	9	8	6331.399			
16	9	7	6331.338			
16	10	7	6629.985		8174.291	
16	10	6	6630.093		8174.284	
16	11	6	6951.509	d		
16	11	5	6951.509	d		
16	12	5	7294.505	d	8881.623	d
16	12	4	7294.505	d	8881.623	d
16	13	4	7657.897	d	9268.805	
16	13	3	7657.897	d	9268.725	
16	14	3	8042.931	d	9670.878	d
16	14	2	8042.931	d	9670.878	d
16	15	2	8446.768	d	10092.855	d
16	15	1	8446.768	d	10092.855	d
16	16	1	8870.284	d		
16	16	0	8870.284	d		
17	0	17	4823.826	d	6180.624	d
17	1	17	4823.826	d	6180.624	d
17	1	16	5078.784		6454.221	
17	2	16	5079.003		6454.638	
17	2	15	5293.768		6682.573	
17	3	15	5297.188		6688.802	
17	3	14	5459.125		6851.657	
17	4	14	5485.970		6892.335	
17	4	13	5572.703		6970.837	
17	5	13	5663.264		7087.895	
17	5	12	5688.369		7107.258	
17	6	12	5853.563			
17	6	11	5856.961			
17	7	11	6071.108			
17	7	10	6071.366			
17	8	10	6317.517			
17	8	9	6317.555			
17	9	9	6590.166			
17	9	8	6590.521			
17	10	8	6887.482			

TABLE 3—Continued

<i>J</i>	<i>K_a</i>	<i>K_c</i>	(020)	(030)
17	10	7	6887.748	
17	11	7	7207.185	8773.701
17	11	6	7207.183	8773.684
17	12	6	7548.568	d
17	12	5	7548.568	d
17	13	5	7909.580	d
17	13	4	7909.580	d
17	14	4	8292.651	d
17	14	3	8292.651	d
17	15	3	8693.918	d
17	15	2	8693.918	d
17	16	2	9114.763	d
17	16	1	9114.763	d
17	17	1	9553.977	d
17	17	0	9553.977	d
18	0	18	5054.907	d 6410.733 d
18	1	18	5054.907	d 6410.733 d
18	1	17	5325.981	6702.406
18	2	17	5326.073	6702.645
18	2	16	5556.514	6947.844
18	3	16	5558.617	6951.889
18	3	15	5740.213	7136.553
18	4	15	5759.595	7167.029
18	4	14	5867.250	7266.800
18	5	14	5942.614	
18	5	13	5980.165	
18	6	13	6133.408	
18	6	12	6139.572	
18	7	12	6349.142	
18	7	11	6349.565	
18	8	11	6593.183	
18	8	10	6593.387	
18	9	10	6864.234	
18	9	9	6864.827	
18	10	9	7159.668	
18	10	8	7159.394	
18	11	7	7477.528	
18	12	7	7816.054	d
18	12	6	7816.055	d
18	13	6	8175.745	d
18	13	5	8175.745	d
18	14	5	8556.839	d
18	14	4	8556.839	d
18	15	4	8955.350	d
18	15	3	8955.350	d
18	16	3	9373.426	d
18	16	2	9373.426	d
18	17	2	9809.778	d
18	17	1	9809.778	d
18	18	1	10258.304	d
18	18	0	10258.304	d
19	0	19	5298.092	d 6652.504
19	1	19	5298.092	d 6652.646
19	1	18	5585.278	6962.716
19	2	18	5585.329	6962.827
19	2	17	5830.849	7224.509
19	3	17	5832.122	7227.091
19	3	16	6032.024	7432.226
19	4	16	6044.858	7454.531
19	4	15	6176.016	7576.713
19	5	15	6235.871	

TABLE 3—Continued

<i>J</i>	<i>K_a</i>	<i>K_c</i>	(020)	(030)
19	5	14	6288.926	
19	6	14	6428.113	
19	6	13	6438.720	
19	7	13	6641.838	
19	7	12	6643.079	
19	8	12	6884.063	
19	8	11	6884.128	
19	9	11	7152.897	
19	9	10	7152.522	
19	10	10	7446.587	
19	10	9	7446.117	
19	11	9	7762.375	
19	11	8	7762.359	
19	12	8	8099.137	
19	12	7	8099.137	a
19	13	7	8456.368	d
19	13	6	8456.368	d
20	0	20	5553.322	d 6906.757
20	1	20	5553.322	d 6906.824
20	1	19	5856.631	7235.211
20	2	19	5856.645	7235.160
20	2	18	6116.857	7512.665
20	3	18	6117.619	7514.248
20	3	17	6334.436	7738.131
20	4	17	6342.986	7754.482
20	4	16	6496.337	7899.724
20	5	16	6542.561	
20	5	15	6613.810	
20	6	15	6737.434	
20	6	14	6754.659	
20	7	14	6950.001	
20	7	13	6951.936	
20	8	13	7189.485	
20	9	12	7456.101	
20	9	11	7455.924	
21	0	21	5820.577	7172.861
21	1	21	5820.616	7172.929
21	1	20	6140.014	7519.627
21	2	20	6140.037	
21	2	19	6414.528	7812.473
21	3	19	6414.949	7813.166
21	3	18	6647.595	
21	4	18	6653.128	
21	4	17	6827.262	
21	5	17	6862.266	
21	6	16	7060.912	
21	6	15	7087.217	
21	7	15	7272.515	
21	7	14	7276.268	
21	8	14	7508.776	
21	8	13	7510.086	
21	10	12	8064.452	
21	10	11	8063.861	
22	0	22	6099.713	7450.882
22	1	22	6099.802	7451.331
22	1	21	6435.418	7816.239
22	2	21	6435.483	7815.555
22	2	20	6723.891	
22	3	20	6724.127	
22	3	19	6971.579	
22	4	19	6975.122	

TABLE 3—Continued

J	K_a	K_c	(020)	(030)
22	4	18	7169.329	
22	5	18	7194.425	
22	5	17	7308.237	
22	6	17	7398.144	
22	6	16	7436.726	
23	0	23	6390.891	7740.686
23	1	23	6390.998	7740.803
23	1	22	6742.804	
23	2	22	6742.966	
23	2	21	7044.892	
23	3	21	7044.949	
23	3	20	7306.531	
23	4	20	7308.736	
24	0	24	6693.846	8042.550 d
24	1	24	6693.943	8042.658 d
24	1	23	7062.126	
24	2	23	7063.575	
24	2	22	7377.377	
24	3	22	7377.530	
24	3	21	7652.538	
24	4	21	7653.801	
24	4	20	7881.869	
25	0	25	7008.773	
25	1	25	7008.797	
25	1	24	7392.866	
25	2	24	7393.400	
25	2	23	7721.393	
25	3	23	7721.804	
25	3	22	8009.664	
26	0	26	7335.478	
26	1	26	7335.497	
26	1	25	7736.133	
26	2	25	7736.662	
27	0	27	7673.981	
27	1	27	7674.006	

TABLE 4—Continued

J	K_a	K_c	E / cm ⁻¹
5	1	5	5653.5189
5	1	4	5703.6235
5	2	4	5758.3228
5	2	3	5768.6395
6	0	6	5727.4602
6	1	6	5734.4833
6	1	5	5803.5002
6	2	5	5849.5472
7	0	7	5823.6422
7	1	7	5828.1982
7	1	6	5918.0485
7	2	5	5985.9790
7	3	5	6149.7993
7	3	4	6153.3642
7	4	4	6320.5506
8	0	8	5931.5867
8	1	8	5934.4373
8	1	7	6046.2411
8	2	7	6074.5291
8	2	6	6120.4486
9	0	9	6051.3254
9	1	9	6053.0712
9	2	8	6207.5327
9	2	7	6271.1832
9	3	7	6417.2246
9	0	9	6051.3254
9	1	9	6053.0712
9	2	8	6207.5327
9	2	7	6271.1832
9	3	7	6417.2246
10	0	10	6182.9611
10	1	10	6184.0147
10	1	9	6339.7061
10	2	8	6504.9100
10	4	7	6747.7909
10	4	6	6749.9688
11	1	11	6327.1599
11	1	10	6503.6295
11	2	10	6583.8419
11	4	8	6921.7836
11	1	11	6327.1599
11	1	10	6503.6295
11	2	10	6583.8419
11	4	8	6921.7836
12	0	12	6482.0818
12	1	12	6482.4608
12	0	12	6482.0818
12	1	12	6482.4608
13	0	13	6649.6576
13	1	13	6649.8965

TABLE 4
Experimental Term Values for the (040) Vibrational
State of HDO

J	K_a	K_c	E / cm ⁻¹
1	0	1	5435.6068
1	1	0	5465.0623
2	0	2	5466.4376
2	1	1	5499.5585
3	0	3	5511.8399
3	1	3	5530.7079
3	1	2	5551.0726
3	2	2	5620.0923
3	2	1	5621.7228
3	3	1	5802.9254
3	3	0	5802.9551
4	0	4	5571.0447
4	1	4	5585.5093
4	2	3	5681.7361
4	2	2	5686.4474
5	0	5	5643.1421

4. CONCLUSIONS

We present the first hot emission spectrum of HDO. This spectrum is very dense since it contains many lines belong to H₂O and D₂O as well as HDO. HDO lines have been identified by comparison with other hot emission spectra. The HDO

transitions have been analyzed resulting in the assignment of many hot pure rotational transitions, bending hot bands linking states up to the previously unobserved (040) bending state, and vibrational difference bands.

ACKNOWLEDGMENTS

We thank A. Rudolph for assistance with the measurements. This work was supported by the Natural Sciences and Engineering Research Council of Canada (NSERC). Some support was also provided by the NASA Laboratory Astrophysics Program, the UK Engineering and Science Research Council, and the Royal Society and the Russian Fund for Fundamental Studies under Grants N 01-02-06411 and N 00-02-16604.

REFERENCES

1. J. Tennyson, N. F. Zobov, R. Williamson, O. L. Polyansky, and P. F. Bernath, *J. Phys. Chem. Ref. Data*, in press.
2. M. Carleer, A. Jenouvrier, A.-C. Vandaele, P. F. Bernath, M. F. Mérienne, R. Colin, N. F. Zobov, O. L. Polyansky, J. Tennyson, and V. A. Savin, *J. Chem. Phys.* **111**, 2444–2450 (1999).
3. N. F. Zobov, D. Belmiloud, O. L. Polyansky, J. Tennyson, S. V. Shirin, M. Carleer, A.-C. Vandaele, P. F. Bernath, M. F. Mérienne, and R. Colin, *J. Chem. Phys.* **113**, 1546–1552 (2000).
4. O. L. Polyansky, N. F. Zobov, S. Viti, J. Tennyson, P. F. Bernath, and L. Wallace, *J. Mol. Spectrosc.* **186**, 422–447 (1997).
5. N. F. Zobov, O. L. Polyansky, J. Tennyson, J. A. Lotoski, P. Colarusso, K.-Q. Zhang, and P. F. Bernath, *J. Mol. Spectrosc.* **193**, 118–136 (1999).
6. N. F. Zobov, O. L. Polyansky, J. Tennyson, S. V. Shirin, R. Nassar, T. Hirao, T. Imajo, P. F. Bernath, and L. Wallace, *Astrophys. J.* **530**, 994–998 (2000).
7. R. Lanquetin, L. H. Coudert, and C. Camy-Peyret, *J. Mol. Spectrosc.* **206**, 83–103 (2001).
8. L. Wallace, P. Bernath, W. Livingston, K. Hinkle, J. Busler, B. Guo, and K.-Q. Zhang, *Science* **268**, 1155–1158 (1995).
9. H. Partridge and D. W. Schwenke, *J. Chem. Phys.* **106**, 4618–4639 (1997).
10. O. L. Polyansky, N. F. Zobov, S. Viti, J. Tennyson, P. F. Bernath, and L. Wallace, *Science* **277**, 348–356 (1997).
11. F. W. Irion, E. J. Moyer, M. R. Gunson, C. P. Rinsland, Y. L. Yung, H. A. Michelsen, R. J. Salawitch, A. Y. Chang, M. J. Newchurch, M. M. Abbas, M. C. Abrams, and R. Zander, *Geophys. Res. Lett.* **23**, 2381–2384 (1996).
12. R. I. Epstein, J. M. Latimer, and D. N. Schramm, *Nature* **263**, 198–207 (1976).
13. R. Meier, T. C. Owen, H. E. Mathews, D. C. Jewitt, D. Bockelée-Morvan, N. Biver, J. Crovisier, and D. Gautier, *Science* **279**, 842–844 (1998).
14. D. W. Schwenke, *J. Phys. Chem. A* **105**, 2352–2360 (2001).
15. R. L. Vander Wal, J. L. Scott, F. F. Crim, K. Weide, and R. Schinke, *J. Chem. Phys.* **94**, 3548–3555 (1991).
16. W. S. Benedict, N. Gailar, and E. K. Plyler, *J. Chem. Phys.* **24**, 1139–1165 (1956).
17. N. Papineau, C. Camy-Peyret, J.-M. Flaud, and G. Guelachvili, *J. Mol. Spectrosc.* **92**, 451–468 (1982).
18. R. A. Toth, *J. Mol. Spectrosc.* **195**, 73–97 (1999).
19. R. A. Toth, *J. Mol. Spectrosc.* **162**, 20–40 (1993).
20. R. A. Toth, V. D. Gupta, and J. W. Brault, *Appl. Opt.* **21**, 3337–3347 (1982).
21. R. A. Toth and J. W. Brault, *Appl. Opt.* **22**, 908–926 (1983).
22. R. A. Toth, *J. Mol. Spectrosc.* **186**, 276–292 (1997).
23. R. A. Toth, *J. Mol. Spectrosc.* **186**, 66–89 (1997).
24. (R. S. Mulliken), *J. Chem. Phys.* **23**, 1997–2011 (1955).
25. A. D. Bykov, Yu. S. Makushkin, V. I. Serdyukov, L. N. Sinita, O. N. Ulenikov, and G. A. Ushakova, *J. Mol. Spectrosc.* **105**, 397–409 (1984).
26. A. D. Bykov, V. P. Lopasov, Yu. S. Makushkin, L. N. Sinita, and V. E. Zuev, *J. Mol. Spectrosc.* **94**, 1–27 (1982).
27. S.-M. Hu, O. N. Ulenikov, G. A. Onopenko, E. S. Bekhtereva, S.-G. He, X.-H. Wang, and Q.-S. Zhu, *J. Mol. Spectrosc.* **203**, 228–234 (2000).
28. J. K. Messer, F. C. DeLucia, and P. Helminger, *J. Mol. Spectrosc.* **105**, 139–155 (1984).
29. O. I. Baskakov, V. A. Alekseev, E. A. Alekseev, and B. I. Pelevoi, *Opt. Spectrosc.* **63**, 600–601 (1987).
30. J. W. C. Johns, *J. Opt. Soc. Am. B* **2**, 1340–1354 (1985).
31. R. Paso and V.-M. Horneman, *J. Opt. Soc. Am. B* **12**, 1813–1838 (1995).
32. O. Naumenko, E. Bertseva, and A. Campargue, *J. Mol. Spectrosc.* **197**, 122–132 (1999).
33. O. Naumenko and A. Campargue, *J. Mol. Spectrosc.* **199**, 59–72 (2000).
34. E. Bertseva, O. Naumenko, and A. Campargue, *J. Mol. Spectrosc.* **203**, 28–36 (2000).
35. A. Campargue, E. Bertseva, and O. Naumenko, *J. Mol. Spectrosc.* **204**, 94–105 (2000).
36. O. Naumenko, E. Bertseva, A. Campargue, and D. W. Schwenke, *J. Mol. Spectrosc.* **201**, 297–309 (2000).
37. A. Jenouvrier, M. F. Mérienne, M. Carleer, R. Colin, A.-C. Vandaele, P. F. Bernath, O. L. Polyansky, and J. Tennyson, *J. Mol. Spectrosc.*, in press.
38. S. Hu, H. Lin, S. He, J. Cheng, and Q. Zhu, *Phys. Chem. Chem. Phys.* **1**, 3727–3730 (1999).
39. A. D. Bykov, V. A. Kapitanov, O. V. Naumenko, T. M. Petrova, V. I. Serdyukov, and L. N. Sinita, *J. Mol. Spectrosc.* **153**, 197–207 (1992).
40. O. L. Polyansky, N. F. Zobov, J. Tennyson, J. A. Lotoski, and P. F. Bernath, *J. Mol. Spectrosc.* **184**, 35–50 (1997).
41. R. A. Toth, *J. Mol. Spectrosc.* **194**, 28–42 (1999).
42. R. A. Toth, *J. Mol. Spectrosc.* **190**, 379–396 (1998).
43. D. W. Schwenke available at <http://george.arc.nasa.gov/~dschwenke/>.
44. O. L. Polyansky, J. Tennyson, and P. F. Bernath, *J. Mol. Spectrosc.* **186**, 213–221 (1997).