# Fourier Transform Emission Spectroscopy of the $A^2\Delta - X^2\Pi$ Transition of SiH and SiD

R. S. Ram, R. Engleman Jr., and P. F. Bernath<sup>2</sup>

Department of Chemistry, University of Arizona, Tucson, Arizona 85721

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The emission spectra of the  $A^2\Delta - X^2\Pi$  transition of SiH and SiD have been observed at high resolution using a Fourier transform spectrometer. The molecules were excited in a Si hollow cathode lamp by passing a discharge through a mixture of Ne and a trace of  $H_2$  or  $D_2$ . The present data, combined with the previous infrared vibration—rotation measurements, have been used to determine improved molecular constants for the ground and excited states of SiH and SiD. © 1998 Academic Press

## INTRODUCTION

SiH is a free radical of fundamental importance. In recent years, there have been numerous experimental and theoretical studies of SiH because of its importance in astrophysics (I) and chemical vapor deposition of thin films (2–4). Because of significant cosmic abundances of Si and H, there is a strong possibility that the SiH radical may be found in the interstellar medium and stellar atmospheres (5). SiH has already been identified in the spectra of sunspots (6–8). As a chemical intermediate this radical plays an important role in many industrial processes such as plasma vapor deposition, thin film formation, and semiconductor manufacturing. The primary diagnostic of silane plasmas has been the observation of the  $A^2\Delta - X^2\Pi$  transition of SiH by optical emission spectroscopy (9–13), laser spectroscopy (14, 15), and laser optogalvanic spectroscopy (16).

The emission spectrum of SiH has been known since 1930 when Jackson (17) observed a strong transition of SiH near 410 nm using an arc source. He obtained a rotational analysis of the 0–0 band and assigned it as the  $A^2\Delta$ – $X^2\Pi$  transition. This analysis was later modified by Mulliken (18). The spectra of SiH and SiD were reinvestigated by Rochester (19), Douglas (20), and Klynning and Lindgren (21). Rochester (19) obtained an analysis of the 0–0 and 1–1 bands of SiH and the 0–0 band of SiD, whereas Douglas (20) observed the SiH bands at a higher dispersion and analyzed 0–0, 1–0, 2–1, and 2–2 bands. In a more recent study Klynning and Lindgren (21) provided an analysis for 1–1 band of SiH and the 0–0, 1–1, 2–2, 1–0, and 2–1 bands of SiD. Several higher-lying excited electronic states,  $B^2\Sigma^+$ ,  $C^2\Sigma^+$ ,  $D^2\Delta$ , and  $E^2\Sigma^+$ , have also been observed for SiH and SiD (22–24). The  $B^2\Sigma^+$  and  $D^2\Delta$  states were found

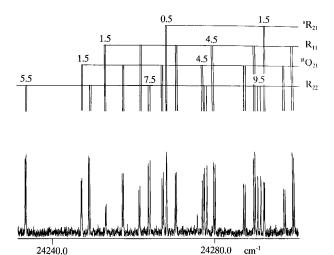
to be strongly predissociated (22, 23). More recently, a new electronic state of SiH and SiD was detected near 46 700 cm<sup>-1</sup> by resonance-enhanced multiphoton ionization spectroscopy (25). This state was tentatively assigned as a  $^2\Pi$  state based on a computer simulation of the spectrum.

Because of astrophysical interest in SiH, transitions in other spectral regions have also been investigated extensively. The lowest pure rotational transition lies outside of the spectral range available to radio astronomers, but lambda doubling transitions can be used to detect SiH. Similar radiofrequency transitions have been successfully used for the detection of OH and CH in the interstellar medium (26, 27). Wilson and Richards (28) obtained a Λ-doubling frequency of 2940  $\pm$  300 MHz for J = 0.5 by extrapolating the  $\Lambda$ -doubling from high rotational levels populated in the optical experiment of Douglas and Elliott (5). Klynning et al. (29) remeasured their high dispersion plates in order to obtain the more precise values of the  $\Lambda$ -doubling constants for the ground  $X^2\Pi$  state. From these measurements a value of 2968  $\pm$  6 MHz was predicted for the  $\Lambda$ -doubling transition. Cooper and Richards (30) also carried out an ab initio calculation of this frequency. Freedman and Irwin (31) have refitted the line positions of the 0-0 and 1-0 bands of Douglas (20) without first deriving term values [direct approach of Zare et al. (32)] and provided a value of 2932 MHz.

The SiH vibration–rotation bands were observed by Knights *et al.* (*33*) in an emission study of a silane plasma in the 1800–2300 cm<sup>-1</sup> region. From this study the vibrational and rotational temperatures of 2000 and 485 K were obtained. The infrared spectra of SiH were also observed by Brown and Robinson (*34*), Brown *et al.* (*35*), Davies *et al.* (*36*), and Betrencourt *et al.* (*37*). Brown and Robinson (*34*) detected the fundamental 1–0 band of SiH by laser magnetic resonance spectroscopy, while Betrencourt *et al.* (*37*) observed the 1–0, 2–1, and 3–2 vibration–rotation bands at a resolution of 0.005

<sup>&</sup>lt;sup>1</sup> Present address: Department of Chemistry, University of New Mexico, Albuquerque, NM 87131.

<sup>&</sup>lt;sup>2</sup> Also at Department of Chemistry University of Waterloo, Waterloo, Ontario, Canada N2L 3G1.



**FIG 1.** A portion of the 0-0 band showing some low J lines in  $R_{11}$ ,  $R_{22}$ ,  ${}^{R}Q_{12}$ , and  ${}^{s}R_{21}$  branches of SiH.

cm<sup>-1</sup> using a Fourier transform spectrometer and provided much improved constants for the ground state. In the most recent study of this radical, Seebass *et al.* (38) have measured several vibration–rotation lines in the 1–0, 2–1, 3–2, 4–3, and 5–4 bands of five isotopomers of SiH and have determined a very precise set of ground state molecular constants including the spin-orbit constant of  $A_0 = 142.8957$  (8) cm<sup>-1</sup>.

The intensity of the  $A^2\Delta - X^2\Pi$  electronic transition of SiH is of concern to astronomers and to material scientists. The absorption cross section and electronic transition moment of the  $A^2\Delta - X^2\Pi$ transition have been determined by Park (39) using a shock tube absorption technique, and a transition moment of  $0.12 \pm 0.04$  au has been obtained. The  $A^2\Delta - X^2\Pi$  transition of SiH has also been observed in vacuum UV photolysis experiments of silane (SiH<sub>4</sub>) and dichlorosilane (SiH<sub>2</sub>Cl<sub>2</sub>) by Washida et al. (40). In a recent paper Stamou et al. (41) have simulated the rotational intensity distribution of the 0–0 band of the  $A^2\Delta - X^2\Pi$  transition of SiH observed in emission from a radio frequency discharge and proposed a new set of molecular constants. There are several studies of the radiative lifetimes of the  $A^2\Delta$  state of SiH (7, 8, 42–45) using the solar spectrum (7, 8, 41), electron bombardment (43-45), and laser-induced fluorescence (14, 46). The solar values, ranging from 569 to 3200 ns, have been estimated from the oscillator strengths  $f_{00}$  and are less reliable. The other experimental values range from 518 to 700 ns. A recent value of 534  $\pm$  23 ns has been obtained using the laser-induced fluorescence technique (46). There are several calculations providing the estimates of dissociation energy of SiH between 3.0 to 3.35 eV. An upper limit of 3.06 eV has been obtained from the analysis of predissociation in the  ${}^{2}\Sigma^{+}$  state (22), while a value of 3.34 eV was obtained from an analysis of the fluorescence lifetimes of the v =1 rotational levels of the  $A^2\Delta$  state (45). A summary of the experimental and theoretical heats of formation of SiH are tabulated by Jasinski et al. (2).

There are several theoretical predictions of the spectroscopic properties of the SiH radical (47-54). A dipole moment of 0.118 D was predicted by Petterson and Langhoff (47), in excellent agreement with the calculated values of Lewerenz *et al.* (48) (0.124 D) and Meyer and Rosmus (49) (0.115 D). Park and Sun (50) have applied the valence shell Hamiltonian method base on quasi-degenerate manybody perturbation theory to calculate the molecular properties including the valence state energies. Theoretical calculations have also been focused on the study of geometrical structures, force constants, and vibrational spectra and heats of formation of silane (SiH<sub>n</sub>) and chlorinated silane molecules (SiH<sub>n</sub>Cl<sub>m</sub>) (51-55).

# **EXPERIMENTAL**

The SiH and SiD molecules were made in a demountable silicon hollow cathode lamp (56). The cathode was prepared by inserting a solid rod of boron-doped silicon into a hole in a copper block. The central part of the rod was then bored through to provide a uniform layer of silicon inside the cathode. The addition of 0.02% of boron to the silicon provided a material with good electrical conductivity. The lamp was operated at 470 V and 610 mA current with a slow flow of neon. Strong SiH bands were initially observed using neon carrier gas at 2 Torr without any added hydrogen. The hydrogen apparently came as impurity in the neon gas or in the vacuum system. The addition of more hydrogen did not increase the intensity of SiH bands. The SiD bands were observed by adding a trace of D<sub>2</sub> gas to the neon flow.

The spectra were recorded using the 1-m Fourier transform spectrometer associated with the McMath-Pierce Solar Telescope of the National Solar Observatory. The spectra in the 17 000–35 000 cm<sup>-1</sup> region were recorded using Si photodiode detectors and a CuSO<sub>4</sub> filter. The SiH bands were recorded at 0.016 cm<sup>-1</sup> resolution after coadding five scans in about 75 min of integration, while the SiD bands were recorded at 0.05 cm<sup>-1</sup> after coadding 25 scans in 130 min of integration.

The spectral line positions were extracted from the observed spectra using data reduction programs called PC-DECOMP and GREMLIN developed by J. Brault. The peak positions were determined by fitting a Voigt line shape function to each spectral feature.

In addition to the SiH/SiD bands, the final spectra also contained Si and Ne atomic lines. The spectra were calibrated using the measurements of Ne atomic lines made by Palmer and Engleman (57). The SiH/SiD lines have widths of about  $0.103 \pm 0.005$  cm<sup>-1</sup> and appear with a maximum signal-tonoise ratio of about 30:1 in the 0–0 band. Because of large widths and weak intensity, the high-J and weaker lines appear to be diffuse and slightly distorted. Part of the strange appearance of the lines could be due to the presence of unresolved

TABLE 1 Vacuum Wavenumbers (in cm  $^{-1}$ ) of the  $A^2\Delta$ - $X^2$ II System of SiH

					SiH	I 0-0	···				
Line	Obs.	о-с	Line	Obs.	О-С	Line	Obs.	о-с	Line	Obs.	о-с
$R_{11e}(1.5)$	24253.164	-4	6.5	24111.326	-5	<sup>S</sup> R <sub>21f</sub> (0.5)	24267.965	0	4.5	24217.099	2
2.5	24261.477	0	7.5	24092.832		1.5	24292.016	3	5.5	24233.569	-1
3.5	24270.503	-1	8.5	24074.515	-2	2.5	24315.827	7	6.5	24249.177	-0
4.5	24280.014	2	9.5	24056.311	-2	3.5	24339.816	1	7.5	24263.974	-2
5.5	24289.836	3	10.5	24038.147	2	4.5	24364.025	8	8.5	24277.995	-3
6.5	24299.838	4	11.5	24019.945		5.5	24388.355	-0	9.5	24291.264	-0
7.5	24309.897		12.5	24001.678	-3	6.5	24412.733	1	10.5	24303.775	-4
8.5	24319.919	1	13.5	23983.274	3	7.5	24437.052	9	11.5	24315.527	
9.5	24329.797	0	14.5	23964.671	0	8.5	24461.183	-2	12.5	24326.494	
10.5	24339.448	1	15.5	23945.819		$^{Q}P_{21e}(3.5)$	24208.526	-6	13.5	24336.661	2
11.5	24348.789	6	16.5	23926.664		4.5	24204.302	0	$R_{22f}(1.5)$	24161.498	
12.5	24357.727	2	$Q_{11e}(2.5)$	24218.338		5.5	24200.138	3	2.5	24181.185	
13.5	24366.189		3.5	24212.625		6.5	24196.073	4	3.5	24199.614	
14.5	24374.114	8	4.5	24207.566		7.5	24192.063	-1	4.5	24216.969	
15.5	24381.386		5.5	24202.936		8.5	24188.053	-6	5.5	24233.350	
16.5	24387.950		6.5	24198.589		9.5	24183.986	-5	6.5	24248.849	
17.5	24393.714		7.5	24194.406		10.5	24179.780		7.5	24263.521	0
18.5	24398.582		8.5	24190.286		11.5	24175.408	6	8.5	24277.388	
$R_{11}(1.5)$	24252.971		9.5	24186.143	-1	12.5	24170.742		9.5	24290.485	
2.5	24261.208		10.5	24181.899		$^{Q}P_{21f}(3.5)$	24208.205	2	10.5	24302.808	
3.5	24270.177		11.5	24177.493		4.5	24203.936	-0	11.5	24314.354	
4.5	24279.649		12.5	24172.830		5.5	24199.750	-7	12.5	24325.105	
5.5	24289.460		13.5	24167.869		6.5	24195.700	-2	13.5	24335.025	
6.5	24299.470		14.5	24162.532		7.5	24191.733	3	14.5	24344.076	
7.5	24309.568		15.5	24156.754		8.5	24187.788	5	$P_{22e}(3.5)$	24068.071	-1
8.5	24319.641		16.5	24150.466		9.5	24183.780	-10	4.5	24057.018	
9.5	24329.598		$Q_{11f}(2.5)$	24218.075	-0	10.5	24179.690	0	5.5	24044.971	-0
10.5	24339.346		3.5	24212.304	7	11.5	24175.408	-6	6.5	24032.148	
11.5	24348.789		4.5	24207.201	-2	12.5	24170.903	4	7.5	24018.664	
12.5	24357.874		5.5	24202.562		$^{R}Q_{21e}(1.5)$	24247.219	-0	8.5	24004.599	
13.5	24366.488		6.5	24198.224	0	2.5	24257.381	-2	9.5	23989.999	
14.5	24374.564		7.5	24194.074	1	3.5	24267.240	1	10.5	23974.905	
15.5	24382.026		8.5	24190.009	1	4.5	24277.212	5	11.5	23959.335	-2
$P_{11e}(2.5)$	24212.396		9.5	24185.942	-1	5.5	24287.313	2	12.5	23943.304	
3.5	24169.502		10.5	24181.791	-7	6.5	24297.496	4	13.5	23926.791	-2
4.5	24149.691		11.5	24177.493	-7	7.5	24307.676	5	14.5	23909.803	4
5.5	24130.498		12.5	24172.978	0	8.5	24317.763	-2	15.5	23892.296	
6.5	24111.706		13.5	24168.165	0	9.5	24327.690	2	$P_{22f}(3.5)$	24067.999	
7.5	24093.160		14.5	24162.994	-1	10.5	24337.355	-7	4.5	24056.881	-3
8.5	24074.798		15.5	24157.399	0	$^{R}Q_{21f}(1.5)$	24247.027	-1	5.5	24044.753	-1
9.5	24056.514		16.5	24151.311	4	2.5	24257.115	2	6.5	24031.817	
10.5	24038.251	2	17.5	24144.654	6	3.5	24266.909	-1	7.5	24018.206	
11.5	24019.945		18.5	24137.347	2	4.5	24276.843	1	8.5	24003.988	-1
12.5 13.5	24001.530		19.5 Sp. (0.5)	24129.317 24268.061	-1 -4	5.5	24286.937	5	9.5 10.5	23989.220	
	23982.970	-3 7	$^{S}R_{21e}(0.5)$			6.5 7.5	24297.128	5		23973.939	
14.5	23964.201		1.5	24292.207 24316.093	2 3	7.5 8.5	24307.343 24317.486	-1	11.5 12.5	23958.167 23941.901	3
15.5	23945.175		2.5		3	8.5	24317.486	-1 -3	12.5	23941.901	-3 -3
16.5	23925.835 23906.112		3.5	24340.147		9.5	24327.483	-3 -2		23923.134 23907.915	-3 5
17.5 D (2.5)			4.5 5.5	24364.382	-0 -5	10.5 11.5	24337.236	-2 -2	14.5 15.5	23890.141	- <b>4</b>
$P_{11f}(2.5)$	24212.123			24388.728 24413.095	-3 -4		24346.713	-2 -1	16.5	23890.141	- <del>4</del> -5
3.5	24169.166		6.5 7.5		-4	R <sub>22e</sub> (1.5) 2.5		-1 -2		24122.509	-3
4.5	24149.324 24130.116		7.5 8.5	24437.375 24461.466		3.5	24181.214 24199.684	-2 -0	$Q_{22e}(2.5)$ 5.5	24122.309	
5.5	Z413U.110	-Z	د.ه	Z <del>44</del> 01.400	4	3.3	24177.004	-0	ر. ر	27132.14/	U

isotope structure for high-J lines. The absolute accuracy and precision of the measurements of sharp and unblended lines is expected to be of the order of  $\pm 0.003~{\rm cm}^{-1}$ . However, the uncertainty of the weaker and blended lines could be as high as  $\pm 0.005~{\rm cm}^{-1}$ .

# RESULTS AND DISCUSSION

The SiH and SiD bands are located in the 21 000–25 000 cm<sup>-1</sup> region that covers the  $\Delta \nu = 0$  and  $\Delta \nu = \pm 1$  sequences. The bands in the  $\Delta \nu = \pm 1$  sequences are very weak and could

TABLE 1—Continued

Line	Obs.	О-С	Line	Obs.	О-С	Line	Obs.	О-С	Line	Obs.	О-С
6.5	24133.575	4	12.5	24126.791	-2	5.5	23975.332	-0	8.5	24006.821	-2
8.5	24134.301	0	13.5	24123.473		6.5	23947.776	-0	9.5	23992.153	1
9.5	24133.695		14.5	24119.471	5	7.5	23919.763	-0	10.5	23977.021	6
10.5	24132.472		15.5	24114.754		8.5	23891.331	-0	11.5	23961.416	-8
11.5	24130.644		${}^{Q}R_{12e}(1.5)$	24122.462	0	9.5	23862.526		12.5	23945.380	1
12.5	24128.194		2.5	24126.588		10.5	23833.365	3	13.5	23928.878	3
13.5	24125.109		3.5	24130.043	-1	$^{0}P_{12f}(4.5)$	24002.271		$^{P}Q_{12f}(2.5)$	24083.436	-6
14.5	24121.357		4.5	24132.722		5.5	23975.117	2	3.5	24072.095	0
16.5	24111.709		5.5	24134.665		6.5	23947.448	-3	4.5	24060.145	-6
17.5	24105.733		6.5	24135.907		7.5	23919.303	-4	5.5	24047.554	-5
$Q_{22f}(2.5)$	24122.478		7.5	24136.486		8.5	23890.713		6.5	24034.346	2
4.5	24129.782			24122.462		9.5	23861.738	-6	7.5	24020.550	-1
5.5	24131.932		2.5	24126.588		10.5	23832.396	1	8.5	24006.219	3
6.5	24133.246		3.5	24129.966		$^{P}Q_{12e}(2.5)$	24083.474	2	9.5	23991.376	1
7.5	24133.817		4.5	24132.596		3.5	24072.165	0	10.5	23976.048	0
8.5	24133.695		5.5	24134.449		4.5	24060.284	1	11.5	23960.250	0
9.5	24132.918		6.5	24135.587		5.5	24047.772	-4	12.5	23943.986	4
10.5	24131.513		7.5	24136.035	-7	6.5	24034.670	-0	13.5	23927.244	5
11.5	24129.469	1	<sup>o</sup> P <sub>12e</sub> (4.5)	24002.408	4	7.5	24021.006	0			
					SiH	1-1					
$R_{11e}(1.5)$	23942.917	-3	12.5	23669.101	-8	$Q_{11f}(2.5)$	23908.779	-6	12.5	24000.741	-19
2.5	23949.661	3	13.5	23645.272	1	3.5	23901.828	-1	$^{R}Q_{21f}(1.5)$	23937.506	-5
3.5	23956.611	5	14.5	23620.644	1	4.5	23895.053	-3	2.5	23945.780	5
4.5	23963.552	9	15.5	23595.140	-2	5.5	23888.253	-1	3.5	23953.382	-3
5.5	23970.311	-0	16.5	23568.670	-10	6.5	23881.278	0	4.5	23960.681	-5
6.5	23976.776	0	17.5	23541.151	-8	7.5	23874.007	-1	5.5	23967.684	1
7.5	23982.811	-3	$P_{11}(2.5)$	23903.572	11	8.5	23866.334	-3	6.5	23974.310	0
8.5	23988.306	-1	3.5	23861.218	1	9.5	23858.165	-2	7.5	23980.476	1
9.5	23993.149	3	4.5	23840.600	3	10.5	23849.400	-3	8.5	23986.086	4
10.5	23997.223		5.5	23820.121	-1	11.5	23839.950	-2	9.5	23991.026	-3
11.5	24000.430		6.5	23799.588	0	12.5	23829.725	2	10.5	23995.214	-1
12.5	24002.676		7.5	23778.871	3	13.5	23818.621	-0	$R_{22e}(2.5)$	23865.748	-12
$R_{11f}(1.5)$	23942.735		8.5	23757.847		14.5	23806.550	-1	3.5	23881.530	-2
2.5	23949.393		9.5	23736.476	5	16.5	23779.091	1	4.5	23895.838	-3
3.5	23956.286		10.5	23714.625	-0	17.5	23763.479	-0	5.5	23908.776	6
4.5	23963.190		11.5	23692.238	-3	$^{S}R_{21e}(1.5)$	23979.912	2	6.5	23920.379	1
5.5	23969.941		12.5	23669.229		2.5	24001.216	3	7.5	23930.696	2
6.5	23976.416		13.5	23645.543	-4	3.5	24022.227	-8	8.5	23939.732	0
7.5	23982.488		14.5	23621.083	3	4.5	24042.974	3	9.5	23947.477	-8
8.5	23988.035		15.5	23595.755	0	5.5	24063.340	-3	10.5	23953.934	1
9.5	23992.942		$Q_{11e}(2.5)$	23909.044	-2	$^{S}R_{21f}(1.5)$	23979.724	-0	11.5	23959.040	1
10.5	23997.118		3.5	23902.148	1	2.5	24000.959	6	12.5	23962.751	-1
11.5	24000.430		4.5	23895.411	1	3.5	24021.923	5	13.5	23965.005	-1
12.5	24002.799		5.5	23888.626	5	4.5	24042.624	7	$R_{22f}(2.5)$	23865.748	15
13.5	24004.083		6.5	23881.637	0	5.5	24062.976	0	3.5	23881.473	6
$P_{11e}(2.5)$	23903.809		7.5	23874.331		$^{R}Q_{21e}(1.5)$	23937.695	0	4.5	23895.715	-2
3.5	23861.547		8.5	23866.613		2.5	23946.033	-3	5.5	23908.569	2
4.5	23840.939		9.5	23858.367		3.5	23953.702	-0	6.5	23920.076	4
5.5	23820.486		10.5	23849.504		4.5	23961.044	5	7.5	23930.262	-2
6.5	23799.943		11.5	23839.950		5.5	23968.049	-1	8.5	23939.152	-5
7.5	23779.201		12.5	23829.590		6.5	23974.672	4	9.5	23946.743	-5
8.5	23758.131		13.5	23818.343	-2	7.5	23980.804	1	10.5	23953.013	-1
9.5	23736.670		14.5	23806.116		8.5	23986.364	8	11.5	23957.926	4
10.5	23714.731		15.5	23792.798		9.5	23991.225	-5	12.5	23961.416	-5
11.5	23692.238	0	16.5	23778.287	-2	11.5	23998.538	4	13.5	23963.446	-0

not be measured. In the  $\Delta \nu=0$  sequence the 0-0 band is the strongest and the 1-1 band is about 20% of the intensity of the 0-0 band. In this analysis we have only included the lines of these two bands. The 2-2 band is very weakly present in our

spectra and a few Q-branch lines were seen near the heads, but the data were not sufficient for a meaningful analysis.

In a  $^2\Delta$ - $^2\Pi$  transition there are six strong main branches,  $P_1$ ,  $P_2$ ,  $Q_1$ ,  $Q_2$ ,  $R_1$ , and  $R_2$ , each of which is doubled by lambda

TABLE 1—Continued

Line	Obs.	о-с	Line	Obs.	О-С	Line	Obs.	о-с	Line	Obs.	о-с
P <sub>22e</sub> (4.5)	23745.374	-2	11.5	23624.373	3	14.5	23756.139	3	4.5	23748.281	2
5.5	23731.540	-6	12.5	23602.902	5	15.5	23743.518	0	5.5	23734.048	-1
6.5	23716.505	-4	13.5	23580.355	1	$Q_{22f}(2.5)$	23810.538	-16	6.5	23718.768	-2
7.5	23700.365	2	14.5	23556.702	4	3.5	23812.939	4	7.5	23702.474	3
8.5	23683.162	-2	15.5	23531.875	2	4.5	23813.794	10	9.5	23666.900	4
9.5	23664.938	-7	$Q_{22e}(1.5)$	23806.147	-3	5.5	23813.270	-4	10.5	23647.634	-3
10.5	23645.717	-3	2.5	23810.577	-5	6.5	23811.498	1	11.5	23627.392	4
12.5	23604.233	4	3.5	23813.000	-1	7.5	23808.511	2	$^{P}Q_{12f}(2.5)$	23773.576	11
13.5	23581.915	1	4.5	23813.914	6	8.5	23804.347	1	3.5	23761.381	1
14.5	23558.496	-5	5.5	23813.476	-2	9.5	23799.017	-3	4.5	23748.157	2
15.5	23533.934	3	6.5	23811.806	3	10.5	23792.528	-2	5.5	23733.847	2
$P_{22}(4.5)$	23745.243	-9	7.5	23808.941	2	11.5	23784.856	1	6.5	23718.466	2
5.5	23731.337	-4	8.5	23804.920	1	12.5	23775.963	-2	7.5	23702.039	-2
6.5	23716.211	7	9.5	23799.759	2	13.5	23765.809	-1	8.5	23684.604	3
7.5	23699.942	8	10.5	23793.445	-3	14.5	23754.334	1	9.5	23666.160	2
8.5	23682.588	-2	11.5	23785.972	-0	15.5	23741.458	-2	10.5	23646.721	3
9.5	23664.203	-4	12.5	23777.298	1	$^{P}Q_{12e}(2.5)$	23773.589	-3			
10.5	23644.795	-6	13.5	23767.370		3.5	23761.444	-0			

doubling in the  ${}^{2}\Pi$  state. These are the only branches that are allowed if both states obey pure Hund's case (a) coupling or if both states have pure Hund's case (b) coupling. For SiH, however, the  $A^2\Delta$  state changes from approximately Hund's case (a) to Hund's case (b) as J increases and several satellite branches become allowed. In total six additional satellite branches ( ${}^{Q}R_{12}$ ,  ${}^{P}Q_{12}$ ,  ${}^{O}P_{12}$ ,  ${}^{S}R_{21}$ ,  ${}^{R}Q_{21}$ , and  ${}^{Q}P_{21}$ ) are expected in each band. In the 0-0 band all of these branches have been identified. In the 1-1 band some of the satellite branches, such as  ${}^{O}P_{12}$  and  ${}^{Q}P_{21}$ , could not be identified because of their very weak intensity. In SiD some of the satellite branches also were not observed in the 0-0 and 1-1 bands. Although Klynning and Lindgren (21) observed more bands in their study, the Fourier transform measurements are expected to be more precise by about an order of magnitude. A part of the 0-0 band of SiH showing some low J lines of a few branches has been provided in Fig. 1. As shown, the  $\Lambda$ -doubling is resolved even for the lowest J lines observed. The observed line positions of SiH and SiD bands have been provided in Tables 1 and 2.

In order to determine the rotational constants, the observed line positions of different bands were fitted with the effective  $N^2$  Hamiltonian of Brown *et al.* (58). The  $^2\Pi$  matrix element of this Hamiltonian is provided by Amiot *et al.* (59), and those for a  $^2\Delta$  state are provided by Brown *et al.* (60). The lines in each of the vibrational bands were initially fitted separately using a nonlinear least-squares procedure. In the final fit the zero field wavenumbers of the 1–0 and 2–1 vibration–rotation bands calculated by Brown (61) (provided for convenience in Table 3), the infrared vibration–rotation bands of Betrencourt *et al.* (37), and far infrared frequencies calculated from the constants of Brown *et al.* (35) provided in the Table 2 of Betrencourt *et al.* (37) were combined with our present measurements. These infrared and far infrared frequencies were weighted by the estimated uncertainties provided in the original papers. The weights for the weaker and blended lines

were chosen according to the signal-to-noise ratio and extent of blending. The constants for SiH and SiD obtained in this fit are provided in Tables 4 and 5. The final reduced standard deviations are 1.20 for SiH and 1.14 for SiD. For convenience, the ground and excited state term values for SiH and SiD are also provided in Tables 6 and 7, respectively. The spectroscopic parameters  $T_{\nu}$ (except v = 0),  $A_v$ ,  $\gamma_v$ ,  $\gamma_{Dv}$  (except v = 0),  $B_v$ ,  $D_v$ ,  $H_v$ ,  $q_v$ ,  $q_{Dv}$ ,  $p_v$ , and  $p_{Dv}$  were determined in the ground state. The term value for the v = 0 vibrational level of the ground state  $(T_0)$  was held fixed to zero. Except for  $q_v$ ,  $q_{Dv}$ ,  $p_v$ , and  $p_{Dv}$ , all the above parameters were determined in the excited  $A^2\Delta$  state. The inclusion of high-J vibration–rotation lines from Betrencourt et al. (37) is expected to improve the determination of distortion constants in the ground state and hence help break the strong correlations with the excited state. The present ground state values are in excellent agreement with the previously reported values (35–38), partly because of the inclusion of all the previous infrared measurements. In particular the present values of  $\Delta G''(1/2) = 1971.03904(35) \text{ cm}^{-1}$ ,  $A_0'' =$ 142.88778(39) cm<sup>-1</sup> can be compared by the recent values of  $\Delta G''(1/2) = 1971.0413 \text{ cm}^{-1}, A_0'' = 142.8952 \text{ cm}^{-1} \text{ obtained by}$ Seebass et al. (38) in their laser magnetic resonance experiment. The values of  $\Delta G''(1/2) = 1970.78 \text{ cm}^{-1}$ ,  $A_0'' = 142.83 \text{ cm}^{-1}$ were obtained by Klynning and Lindgren (21). Some of the discrepancies are caused by our adoption of the  $N^2$  Hamiltonian rather than the energy level expressions used by Klynning and Lindgren (21). It has been shown by Brown and Watson (62) that the  $A_{Dv}$  and  $\gamma_v$  parameters cannot be determined simultaneously for a  ${}^{2}\Pi$  state. In their analysis Klynning and Lindgren (21) determined  $A_{Dv}$  ( $A_I$  in their notation) instead of  $\gamma_v$ . The separation of these two parameters is possible only by isotopic substitution, and Betrencourt et al. (37) decided to determine  $\gamma_{\nu}$  rather than  $A_{D\nu}$ . In the most recent infrared study of this radical, Seebass *et al.* (38) observed the spectra of five isotopic species <sup>28</sup>SiH, <sup>29</sup>SiH, <sup>30</sup>SiH, <sup>28</sup>SiD, and <sup>29</sup>SiD and found an appreciable magnitude for

TABLE 2 Vacuum Wavenumbers (in cm  $^{-1}$ ) of the  $A^2\Delta$ - $X^2$ II System of SiD

					SiE	0-0			<del></del>	<del> </del>	
Line	Obs.	о-с	Line	Obs.	О-С	Line	Obs.	0-С	Line	Obs.	о-с
R <sub>11e</sub> (2.5)	24305.365	3	20.5	24078.383	-14	9.5	24260.038	1	10.5	24255.619	4
3.5	24309.328	3	21.5	24067.621	9	10.5	24257.159	1	11.5	24252.806	3
4.5	24313.503	5	22.5	24056.681	6	11.5	24254.303	4	12.5	24249.974	-1
5.5	24317.812	-3	23.5	24045.564	-0	12.5	24251.438		13.5	24247.098	-13
6.5	24322.236	2	25.5	24022.745	13	13.5	24248.554	3	$^{R}Q_{21e}(3.5)$	24306.531	-5
7.5	24326.729	2	$P_{11f}(6.5)$	24223.628	14	14.5	24245.622		4.5	24311.163	-7
8.5	24331.265		7.5	24212.986	6	15.5	24242.626	-3	5.5	24315.780	4
9.5	24335.826		8.5	24202.466		16.5	24239.544	-8	6.5	24320.390	-0
10.5	24340.381		9.5	24192.054		17.5	24236.364	-4	7.5	24325.025	6
11.5	24344.911		10.5	24181.708		18.5	24233.054		8.5	24329.649	
12.5	24349.386		11.5	24171.422		19.5	24229.602	3	9.5	24334.283	
13.5	24353.798		12.5	24161.173		20.5	24225.969		10.5	24338.889	
14.5	24358.102		13.5	24150.932		21.5	24222.150		11.5	24343.439	
15.5	24362.293		15.5	24130.444		22.5	24218.086		12.5	24347.953	
16.5	24366.345		16.5	24120.161		23.5	24213.810		13.5	24352.357	
17.5	24370.208		17.5	24109.832		24.5	24209.259		14.5	24356.682	
18.5	24373.902		18.5	24099.423		25.5	24204.409		15.5	24360.873	
19.5	24377.356		19.5	24088.925		26.5	24199.223		16.5	24364.921	5
21.5	24383.544		20.5	24078.324		27.5	24193.706		17.5	24368.791	7
22.5	24386.199	13	21.5	24067.603		28.5	24187.795	11	18.5	24372.469	
23.5	24388.513		22.5	24056.728		29.5	24181.444		$^{R}Q_{21f}(2.5)$	24301.604	3
$R_{11}(3.5)$	24309.150		23.5	24045.690		30.5	24174.664		3.5	24306.338	
4.5	24313.270		24.5	24034.448		${}^{S}R_{21e}(4.5)$	24356.842		4.5	24310.944	
5.5	24317.556		25.5	24023.007		$K_{21e}(4.5)$ 5.5	24368.931	6	5.5	24315.525	
6.5	24321.945		$Q_{11e}(2.5)$	24023.007		6.5	24381.008	-5	6.5	24313.323	
7.5	24321.943		$Q_{11e}(2.5)$ 3.5	24279.229		7.5	24393.097		7.5	24324.715	
8.5			4.5	24275.723		8.5	24405.177		8.5	24329.331	3
9.5	24330.947		5.5	24273.723		9.5	24403.177	0	9.5	24329.331	9
10.5	24335.490		6.5	24272.420		10.5	24429.188	0	10.5	24333.933	
11.5	24340.029 24344.573		7.5	24266.238		11.5	24441.098	9	11.5	24336.349	4
12.5	24344.373		7.3 8.5	24263.279		12.5	24441.098	4	12.5	24343.111	
						13.5	24464.554		13.5		
13.5	24353.469		9.5	24260.370 24257.498				-2 2	14.5	24352.051	-12
14.5	24357.802		10.5			$^{S}R_{211}(4.5)$	24356.625 24368.680		15.5	24356.371	
15.5	24362.015		11.5 12.5	24254.643 24251.771		5.5 7.5	24392.786		16.5	24360.585	-13
16.5	24366.083								17.5	24364.657	
17.5	24369.994		13.5	24248.871	-1	8.5	24404.849	2	17.5	24368.570	-5
18.5	24373.729		14.5	24245.923		9.5	24416.867	-5		24372.297	7
19.5	24377.236		15.5	24242.901	-5	10.5	24428.845	-2	$R_{22e}(1.5)$	24184.644	
20.5	24380.522		16.5	24239.792		11.5	24440.748	-0	2.5	24195.803	-8
21.5	24383.521		17.5	24236.578		12.5	24452.546		3.5	24206.362	
22.5	24386.234		18.5	24233.226		13.5	24464.256		4.5	24216.508	-14
23.5	24388.626		19.5	24229.713		14.5	24475.779		5.5	24226.277	-6
$P_{11e}(5.5)$	24234.655		20.5	24226.038		$^{Q}P_{21e}(5.5)$	24270.098	-5	6.5	24235.687	-5
6.5	24223.891		21.5	24222.150		6.5	24267.242	-1	7.5	24244.763	
7.5	24213.295		22.5	24218.057		7.5	24264.398	2	8.5	24253.509	1
8.5	24202.806		23.5	24213.691		8.5	24261.574	3	9.5	24261.927	-3
9.5	24192.396		24.5	24209.082		9.5	24258.773	12	10.5	24270.044	10
10.5	24182.055		25.5	24204.153		10.5	24255.953	-3	11.5	24277.821	2
11.5	24171.758		26.5	24198.892		11.5	24253.127		12.5	24285.295	7
12.5	24161.506		27.5	24193.274		12.5	24250.321	12	13.5	24292.441	5
13.5	24151.241		$Q_{11f}(2.5)$	24282.944	-28	13.5	24247.436	3	14.5	24299.264	4
14.5	24141.001		3.5	24279.042		$^{Q}P_{21f}(4.5)$	24272.710		15.5	24305.749	-3
15.5	24130.723	-3	4.5	24275.510	5	5.5	24269.852	5	16.5	24311.902	-5
16.5	24120.403	-8	5.5	24272.170		6.5	24266.954	-6	17.5	24317.704	-9
17.5	24110.027	-12	6.5	24268.998	1	7.5	24264.089	-1	18.5	24323.152	-7
18.5	24099.585	-7	7.5	24265.934	1	8.5	24261.252	6	19.5	24328.237	4
19.5	24089.045	-5	8.5	24262.956	1	9.5	24258.427	2	20.5	24332.920	3

TABLE 2—Continued

<del></del>				TABL		-Commu					
Line	Obs.	о-с	Line	Obs.	о-с	Line	Obs.	о-с	Line	Obs.	О-С
21.5	24337.210		11.5	24089.884	8	5.5	24173.148		4.5	24135.404	
22.5	24341.059		12.5	24082.717	7	7.5	24176.678		5.5	24129.787	
23.5	24344.454		13.5	24075.320	7	8.5	24177.986		6.5	24123.951	-10
24.5	24347.396		14.5	24067.682	-2	9.5	24179.007	5		24117.897	
25.5	24349.853		15.5	24059.826	-2	10.5	24179.731	1	8.5	24111.615	
$R_{221}(1.5)$	24184.644		16.5	24051.729	-14	11.5	24180.194		9.5	24105.104	
2.5	24195.803		17.5	24043.428	3	14.5	24179.870		10.5	24098.357	
3.5	24206.362		18.5	24034.860	-12	15.5	24179.214			24091.379	
4.5	24216.508		19.5	24026.065	-9	16.5	24178.297	17	12.5	24084.183	
5.5	24226.244		20.5	24017.029	3	17.5	24177.047			24076.760	
6.5	24235.637		21.5	24007.712	-5	18.5	24175.532		14.5	24069.122	
7.5	24244.691		22.5	23998.139	6	19.5	24173.702	-1	15.5	24061.247	
8.5	24253.408		23.5	23988.287	24	20.5	24171.549	-7	16.5	24053.159	
9.5	24261.791		$P_{22f}(6.5)$	24121.852	-19	21.5	24169.081	3	17.5	24044.832	
10.5	24269.852		7.5	24115.986	1	22.5	24166.249		$^{R}Q_{12f}(3.5)$	24140.847	
11.5	24277.595		8.5	24109.811	8	23.5	24163.057	-5	4.5	24135.378	
12.5	24285.009		9.5	24103.352	8	$Q_{22f}(5.5)$	24173.109		5.5	24129.757	
13.5	24292.100		10.5	24096.620	-3	6.5	24175.022		6.5	24123.912	
14.5	24298.866 24305.289		11.5	24089.644	-6 -3	7.5 8.5	24176.612 24177.885	4	7.5	24117.825	
15.5			12.5 13.5	24082.430		13.5	24177.883	-0 1	8.5 9.5	24111.517	
16.5 17.5	24311.370		13.5	24074.965	-12 4	13.5		1 -0	10.5	24104.956	
	24317.090		15.5	24067.290		15.5	24179.473	-0 5		24098.174	
18.5 19.5	24322.455		16.5	24059.364 24051.192	1	16.5	24178.755	8	11.5 12.5	24091.156	11 6
20.5	24327.456 24332.036		17.5	24031.192	-12 -2	17.5	24177.748 24176.444	8	13.5	24083.901 24076.412	-5
21.5	24336.220		18.5	24042.800	-1	18.5	24174.826		14.5	24070.412	16
22.5	24339.977		19.5	24025.278	-9	19.5	24174.820	8	15.5	24060.728	10
24.5	24339.977		20.5	24025.278	-19	20.5	24172.924	3	16.5	24052.625	4
25.5	24348.461	4	21.5	24016.123	-10	22.5	24170.081		17.5	24032.023	-8
$P_{22e}(5.5)$	24127.469		22.5	23997.054	-7	23.5	24161.874		18.5	24035.602	2
6.5	24121.917		23.5	23987.103	14	24.5	24158.212		19.5	24026.726	
7.5	24116.062		24.5	23976.815	5	25.5	24154.140	21	20.5	24017.588	-11
8.5	24109.901	-6	$Q_{22e}(2.5)$	24164.882	-7	26.5	24149.596	-4	20.5	21017.500	
9.5	24103.478		3.5	24168.141		$^{R}Q_{12e}(2.5)$	24146.254	-7			
10.5	24096.799		4.5	24170.847	Õ	3.5	24140.847	-1			
					SiD	1-1				,	
D (2.5)	24103.949	-1	12.5	24040.310	3	7.5	24035.647	6	16.5	23953.831	14
$R_{11e}(2.5)$			13.5	24040.310	6	8.5	24033.047	13		23955.651	
3.5 4.5	24107.222		14.5	24030.286	19		24042.724	11	$Q_{22f}(4.5)$ 5.5	23969.042	2 -5
5.5	24110.557 24113.860		15.5	24030.286	4	$R_{22f}(3.5)$ 4.5	24002.482	9	10.5	23968.692	3
6.5	24117.130		16.5	24019.209	-10	5.5	24020.046	20	11.5	23967.271	8
8.5	24117.130	-1	$Q_{11}(4.5)$	24074.108	10	6.5	24028.046	3	12.5	23965.394	5
9.5	24126.257		5.5	24070.067	-11	7.5	24035.575	5	13.5	23963.048	-14
10.5	24128.995	3	6.5	24066.066	18	8.5	24042.627	14	14.5	23960.281	4
$R_{11f}(2.5)$	24103.829		7.5	24061.978	5	$P_{22e}(5.5)$	23925.281	17	15.5	23957.024	-3
3.5	24107.055	5	8.5	24057.829	6	6.5	23918.961	-8	16.5	23953.299	1
4.5	24110.343	2	9.5	24053.562	-9	7.5	23912.195		$^{R}Q_{12e}(3.5)$	23939.515	-5
6.5	24116.833		10.5	24049.209	12	8.5	23905.017	8	4.5	23933.612	-21
8.5	24123.034		11.5	24044.668	-6	$P_{22f}(5.5)$	23925.245	12	5.5	23927.402	-23
9.5	24125.917		12.5	24039.973	-7	6.5	23918.921	0	6.5	23920.880	22
10.5	24128.650		13.5	24035.099	8	8.5	23904.908	-2	7.5	23913.911	-3
$Q_{11e}(4.5)$	24074.328		14.5	24029.975	-6	$Q_{22e}(4.5)$	23967.698	-16	8.5	23906.577	-10
5.5	24070.340		15.5	24024.638	10	5.5	23969.090	13	9.5	23898.897	23
6.5	24066.336		16.5	24019.000	-3	10.5	23968.859		$^{R}Q_{12f}(3.5)$	23939.515	4
7.5	24062.278	-4	$R_{22e}(2.5)$	23992.866	-4	11.5	23967.469	-9	4.5	23933.612	-3
8.5	24058.141	-7	3.5	24002.482	2	12.5	23965.648	-6	5.5	23927.402	8
9.5	24053.900		4.5	24011.518	-10	13.5	23963.369	-14	6.5	23920.826	16
10.5	24049.537		5.5	24020.046	-11	14.5	23960.659	0	8.5	23906.501	13
11.5	24045.018		6.5	24028.107	16	15.5	23957.474	-0	9.5	23898.752	10
	2.010.010		0.5	2.020.107							

TABLE 3 Hyperfine-Free Vibration-Rotation Lines (in  $\rm cm^{-1}$ ) of SiH and SiD (61)

<b>J</b> ′	Level	J″	Level	Obs.	0-С	J'	Level	J″	Level	Obs.	О-С
					SiH	1-0					
2.5 2.5 2.5 2.5 1.5 3.5 4.5 5.5 6.5 8.5	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	1.5 1.5 2.5 2.5 2.5 4.5 5.5 6.5 7.5 9.5	F <sub>2</sub> e F <sub>2</sub> f F <sub>2</sub> e F <sub>2</sub> e	2008.3766 2008.3985 1969.3560 1969.4170 1931.6058 1897.9489 1880.6228 1862.9860 1845.0475 1808.2864	-0.0015 0.0008 -0.0027 0.0016 -0.0003 0.0000 0.0000 -0.0000 0.0000 0.0012	1.5 3.5 4.5 5.5 6.5 5.5 0.5 1.5	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	2.5 4.5 5.5 6.5 7.5 6.5 1.5 2.5 1.5 2.5	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	1931.582; 1897.8800 1880.528; 1862.866; 1844.902; 1872.335; 1818.769; 1800.061; 1818.8584; 1800.216;	0 -0.0018 9 -0.0001 2 0.0015 2 0.0040 2 -0.0005 5 0.0008 7 0.0010 4 0.0013
					SiH	2-1					
2.5 2.5 1.5 2.5 4.5 2.5	$\begin{array}{cccc} F_2 & e \\ F_2 & f \\ F_2 & e \\ F_2 & e \\ F_2 & e \\ F_2 & f \end{array}$	1.5 1.5 2.5 3.5 5.5 3.5	$\begin{array}{cccc} F_2 & e \\ F_2 & f \\ F_2 & e \\ F_2 & e \\ F_2 & e \\ F_2 & f \end{array}$	1936.2085 1936.2253 1861.8364 1845.6655 1812.2957 1845.6258	0.0003 -0.0011 -0.0000 -0.0007 0.0015 -0.0004	4.5 4.5 4.5 1.5 1.5 3.5	$\begin{array}{cccc} F_2 & f \\ F_1 & e \\ F_1 & f \\ F_2 & f \\ F_2 & e \\ F_1 & e \\ \end{array}$	5.5 5.5 5.5 1.5 1.5 3.5	$\begin{array}{cccc} F_2 & f \\ F_1 & e \\ F_1 & f \\ F_2 & e \\ F_2 & f \\ F_1 & f \end{array}$	1812.2053 1819.8522 1819.8294 1899.6253 1899.6126	2 -0.0008 4 0.0014 5 0.0006 6 0.0012
					SiD	1-0					
1.5 2.5 3.5 1.5 2.5 3.5 2.5 3.5 4.5	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	1.5 2.5 3.5 1.5 2.5 3.5 1.5 2.5 3.5	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	1432.9958 1432.5704 1431.9771 1432.9938 1432.5626 1431.9581 1452.3405 1459.6308 1466.7317	0.0002 0.0003 0.0002 0.0002 0.0002 0.0001 0.0026 0.0012 -0.0028	2.5 3.5 4.5 1.5 2.5 3.5 1.5 2.5 3.5	F <sub>2</sub> f F <sub>2</sub> f F <sub>2</sub> f F <sub>2</sub> e F <sub>2</sub> e F <sub>2</sub> e F <sub>2</sub> f F <sub>2</sub> f	1.5 2.5 3.5 2.5 3.5 4.5 2.5 3.5 4.5	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	1452.3433 1459.6317 1466.7400 1413.2208 1404.9034 1396.4341 1413.2179 1404.8970 1396.4254	7 -0.0032 6 -0.0022 8 -0.0022 4 -0.0010 1 0.0012 9 -0.0022 6 -0.0008
					SiD	2-1					
1.5 2.5 3.5 1.5 2.5 3.5 2.5 3.5 4.5	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	1.5 2.5 3.5 1.5 2.5 3.5 1.5 2.5 3.5	F <sub>2</sub> e F <sub>2</sub> e F <sub>2</sub> e F <sub>2</sub> f F <sub>2</sub> f F <sub>2</sub> f F <sub>2</sub> e F <sub>2</sub> e	1396.0475 1395.6296 1395.0468 1396.0456 1395.6220 1395.0281 1414.9716 1422.1018 1429.0462	0.0014 -0.0007 -0.0010 0.0014 -0.0007 -0.0012 0.0018 0.0001 0.0001	2.5 3.5 4.5 1.5 2.5 3.5 1.5 2.5 3.5	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	1.5 2.5 3.5 2.5 3.5 4.5 2.5 3.5 4.5	F <sub>2</sub> f F <sub>2</sub> f F <sub>2</sub> f F <sub>2</sub> e F <sub>2</sub> e F <sub>2</sub> e F <sub>2</sub> f F <sub>2</sub> f F <sub>2</sub> f	1414.9743 1422.1072 1429.0549 1376.7008 1368.5616 1360.2735 1376.6980 1368.5559 1360.2642	2 -0.0001 9 -0.0003 8 -0.0010 6 -0.0016 5 0.0019 9 -0.0012 9 -0.0019

the  $\gamma_{\nu}$  parameter. Our results are consistent with the results of these two groups.

Since the ground state molecular parameters are known with excellent precision from the previous infrared studies (35-38), the main purpose of the present study was to improve the molecular constants for the excited  $A^2\Delta$  state. It is therefore more interesting to compare the present excited state constants for the  $A^2\Delta$  state with the values obtained previously by Klynning and Lindgren (21) and Klynning *et al.* (29). Our  $\Delta G'(1/2) = 1662.36001(82)$  cm<sup>-1</sup>,  $A'_0 = 3.54137(82)$  cm<sup>-1</sup> are significantly different from the values of  $\Delta G'(1/2) = 1660.58$  cm<sup>-1</sup>,  $A'_0 = 3.62$  cm<sup>-1</sup> obtained by Klynning and Lindgren (21). The rotational constants for the v = 0 vibrational level of the  $A^2\Delta$ 

state from their revised analysis of the 0–0 band [Klynning *et al.* (29)], however, compare better with our values. In this work they determined  $\gamma_0$  instead of  $A_{D0}$  in the excited state while still keeping  $A_{D0}$  in the ground state. Their  $B_0'=7.28280(8)$  cm<sup>-1</sup>,  $D_0'=5.203(3)\times 10^4$  cm<sup>-1</sup>,  $A_0'=3.544(2)$  and  $\gamma_0=0.0871$  cm<sup>-1</sup> compare well with our values of  $B_0'=7.287599(14)$  cm<sup>-1</sup>,  $D_0'=5.1881(13)\times 10^4$  cm<sup>-1</sup>,  $A_0'=3.54137(82)$  and  $\gamma_0=0.089983(99)$  cm<sup>-1</sup>.

The molecular constants of the v=0 and v=1 vibrational levels have been used to determine the equilibrium constants for SiH and SiD from an exact fit. The excited state equilibrium rotational constants obtained from this work are  $B_e=7.502718(21)~{\rm cm}^{-1}$ ,  $\alpha_e=0.215119(15)~{\rm cm}^{-1}$  for SiH and  $B_e=0.215119(15)~{\rm cm}^{-1}$ 

TABLE 4 Spectroscopic Constants (in cm $^{-1}$ ) for the  $A^2\Delta - X^2$ II System of SiH

Constants <sup>a</sup>	$X^2\Pi$ (v=0)	X <sup>2</sup> ∏ (v=1)	$A^2\Delta$ (v=0)	$A^2\Delta$ (v=1)
$T_{ m v}$	0	1971.03904(35)	24171.36819(44)	25833.72820(69)
$A_{v}$	142.88778(39)	143.46046(46)	3.54137(82)	3.0935(11)
$10^2 \times \gamma_v$	-4.5762(54)	-4.2215(89)	8.9983(99)	8.527(13)
$10^5 \times \gamma_{Dv}$			-2.158(49)	-2.382(71)
$\mathrm{B_{v}}$	7.3917453(62)	7.175928(17)	7.287599(14)	6.857362(26)
$10^4 \times D_v$	4.0174(10)	3.9624(19)	5.1881(13)	5.8242(25)
$10^8 \times H_v$	1.545(25)	1.378(50)	-1.335(28)	-4.711(64)
$10^3 \times q_v$	8.3656(80)	8.147(13)		
$10^6 \times q_{Dv}$	-1.648(36)	-1.733(60)		
$10^2 \times p_v$	8.3316(59)	7.980(11)		
$10^5 \times p_{Dv}$	-1.572(44)	-1.548(68)		

<sup>&</sup>lt;sup>a</sup>Numbers in parentheses are one standard deviation in last two digits.

= 3.878272(77) cm<sup>-1</sup>,  $\alpha_e$  = 0.076744(68) cm<sup>-1</sup> for SiD. The values reported in parentheses are estimated uncertainties. Since these values have been determined from exact fits, the actual uncertainties in these values are expected to be much higher than the quoted values. The equilibrium bond lengths for the excited  $A^2\Delta$  state of SiH and SiD using these values are 1.5197816(21) Å and 1.521017(15) Å, respectively.

In a previous study of this molecule, Verma (22) observed strong predissociation in the  $B^2\Sigma^+$  and  $D^2\Delta$  states. From an analysis of these predissociations, Verma (22) estimated an upper limit of the dissociation energy of 3.06 eV for this

molecule. The lifetimes of a number of rotational levels of the v=0,1, and 2 vibrational levels were measured by Carlson *et al.* (45) using the high-frequency deflection technique. They observed some decrease in the lifetimes of rotational levels at  $N' \geq 12$  in v=1  $F_1$  and  $F_2$  components. They found that the rotational levels up to  $N' \approx 11$  of v=1 have average lifetimes of 594  $\pm$  10 ns, and for higher N' this value decreases slowly to 470 ns in  $F_1$  (at N'=18) and 405 ns in  $F_2$  (at N'=16). They also found that the v=2 vibrational level has a very short lifetime of about 160 ns. From this observation, they concluded that the v=1 and 2 vibrational levels of SiH are affected by

TABLE 5 Spectroscopic Constants (in cm $^{-1}$ ) for the  $A^2\Delta$ - $X^2$ II System of SiD

Constants	$X^2\Pi$ (v=0)	$X^2\Pi$ (v=1)	$A^2\Delta$ (v=0)	$A^2\Delta$ (v=1)
$T_{v}$	0	1433.0351(23)	24224.41130(93)	25456.9993(43)
$A_{\nu}$	142.7487(13)	143.1823(43)	3.6159(41)	3.333(14)
$10^2 \times \gamma_v$	-2.832(18)	-3.12(22)	4.725(16)	4.244(89)
$10^6 \times \gamma_{Dv}$			-5.94(39)	-9.4(42)
$\mathbf{B}_{\mathrm{v}}$	3.846656(35)	3.765807(81)	3.801528(37)	3.64804(13)
$10^4 \times D_v$	1.0916(12)	1.0656(86)	1.3769(12)	1.468(11)
$10^9 \times H_v$	2.98(12)	7.7(19)	-1.11(12)	
$10^3 \times q_v$	2.236(12)	2.149(28)		
$10^7 \times q_{Dv}$	-2.29(22)			
$10^2 \times p_v$	4.375(15)	4.436(50)		
$10^5 \times p_{Dv}$	-0.604(38)	-1.85(29)		

<sup>&</sup>lt;sup>a</sup>Numbers in parentheses are one standard deviation in last two digits.

 ${\bf TABLE~6} \\ {\bf Term~Values~(in~cm^{-1})~for~the~Ground~and~Excited~States~of~SiH}$ 

		$A^2$	$\Delta_{3/2}$			$\mathbf{A}^2$	Δ <sub>5/2</sub>	
-	v=(		v=	1	v=(	`		1
J	e	f	e e	f	ee	, f	e	f
1.5	24211.3990	24211.3990	25871.6302	25871.6302				
2.5	24217.3466	24217.3466	25876.8545	25876.8545	24256.3836	24256.3836	25913.8446	25913.8446
3.5	24260.4771	24260.4771	25917.4672	25917.4672	24315.0907	24315.0907	25969.0230	25969.0230
4.5	24318.3566	24318.3566	25971.9264	25971.9264	24387.9959	24387.9959	26037.5553	26037.5553
5.5	24390.8003	24390.8003	26040.0589	26040.0589	24475.1713	24475.1713	26119.4877	26119.4877
6.5 7.5	24477.6936 24578.9365	24477.6936 24578.9365	26121.7486 26216.8886	26121.7486 26216.8886	24576.5941 24692.2014	24576.5941 24692.2014	26214.7807 26323.3564	26214.7807 26323.3564
8.5	24694.4279	24694.4279	26325.3671	26325.3671	24821.9064	24821.9064	26445.1118	26445.1118
9.5	24824.0593	24824.0593	26447.0624	26447.0624	24965.6036	24965.6036	26579.9240	26579.9240
10.5	24967.7125	24967.7125	26581.8403	26581.8403	25123.1721	25123.1721	26727.6519	26727.6519
11.5	25125.2581	25125.2581	26729.5522	26729.5522	25294.4759	25294.4759	26888.1371	26888.1371
12.5	25296.5546	25296.5546	26890.0347	26890.0347	25479.3653	25479.3653	27061.2042	27061.2042
13.5	25481.4481	25481.4481	27063.1090	27063.1090	25677.6764	25677.6764	27246.6605	27246.6605
14.5	25679.7721	25679.7721	27248.5796	27248.5796	25889.2318	25889.2318	27444.2957	27444.2957
15.5	25891.3468	25891.3468	27446.2345	27446.2345	26113.8403	26113.8403	27653.8820	27653.8820
16.5	26115.9794	26115.9794	27655.8442	27655.8442	26351.2966	26351.2966	27875.1731	27875.1731
17.5	26353.4634	26353.4634	27877.1614	27877.1614	26601.3820	26601.3820	28107.9041	28107.9041
18.5	26603.5788	26603.5788	28109.9201	28109.9201	26863.8632	26863.8632	28351.7908	28351.7908
19.5	26866.0917	26866.0917	28353.8353	28353.8353	27138.4933	27138.4933	28606.5291	28606.5291
20.5	27140.7542	27140.7542	28608.6024	28608.6024	27425.0105	27425.0105	28871.7945	28871.7945
21.5	27427.3042	27427.3042	28873.8963	28873.8963	27723.1388	27723.1388	29147.2413	29147.2413
22.5	27725.4650	27725.4650	29149.3706	29149.3706	28032.5875	28032.5875	29432.5019	29432.5019
23.5	28034.9455	28034.9455	29434.6575	29434.6575	28353.0509	28353.0509	29727.1862	29727.1862
24.5	28355.4396	28355.4396	29729.3663	29729.3663	28684.2081	28684.2081	30030.8806	30030.8806
25.5	28686.6260	28686.6260	30033.0833	30033.0833	29025.7229	29025.7229	30343.1475	30343.1475
		X <sup>2</sup> I	I <sub>1/2</sub>			$X^2I$	$I_{3/2}$	_
0.5	-56.6662	-56.5662	1913.6533	1913.7493				
1.5	-35.8208	-35.6286	1933.9353	1934.1202	94.8846	94.8922	2065.4806	2065.4877
2.5	-0.9994	-0.7293	1967.8093	1968.0696	133.8745	133.9039	2103.2626	2103.2899
3.5	47.8518	48.1804	2015.3205	2015.6381	188.3121	188.3821	2156.0227	2156.0879
4.5	110.7888	111.1539	2076.5162	2076.8700	258.0738	258.2061	2223.6471	2223.7709
5.5	187.8607	188.2386	2151.4374	2151.8050	343.0244	343.2418	2306.0098	2306.2138
6.5	279.1021	279.4693	2240.1125	2240.4711	443.0237	443.3491	2402.9783	2403.2845
7.5	384.5300	384.8634	2342.5537	2342.8810	557.9308	558.3862	2514.4177	2514.8475
8.5	504.1416	504.4191	2458.7555	2459.0300	687.6052	688.2118	2640.1926	2640.7664
9.5	637.9152	638.1157	2588.6943	2588.8954	831.9073	832.6848	2780.1669	2780.9041
10.5	785.8107	785.9142	2732.3292	2732.4374	990.6973	991.6644	2934.2037	2935.1225
11.5	947.7708	947.7586	2889.6030	2889.6000	1163.8343	1165.0084	3102.1646	3103.2817
12.5	1123.7228	1123.5769	3060.4434	3060.3118	1351.1751	1352.5724	3283.9081	3285.2395
13.5	1313.5796	1313.2831	3244.7642	3244.4876	1552.5731	1554.2089	3479.2896	3480.8498
14.5	1517.2408	1516.7775	3442.4659	3442.0287	1767.8777	1769.7661	3688.1598	3689.9625
15.5	1734.5935	1733.9485	3653.4370	3652.8247	1996.9335	1999.0876	3910.3646	3912.4224
16.5	1965.5137	1964.6726	3877.5548	3876.7535	2239.5797	2242.0117	4145.7446	4148.0691
17.5	2209.8666	2208.8158	4114.6856	4113.6826	2495.6496	2498.3708	4394.1345	4396.7362
18.5	2467.5071	2466.2340	4364.6858	4363.4690	2764.9710	2767.9916	4655.3629	4658.2514
19.5	2738.2807	2736.7735	4627.4022	4625.9605	3047.3652	3050.6945	4929.2524	4932.4363
20.5	3022.0239	3020.2715	4902.6724	4900.9956	3342.6474	3346.2938	5215.6189	5219.1058
21.5	3318.5645	3316.5567	5190.3252	5188.4039	3650.6266	3654.5974	5514.2725	5518.0689
22.5	3627.7220	3625.4495	5490.1814	5488.0071	3971.1056	3975.4073	5825.0167	5829.1280
23.5	3949.3082	3946.7626	5802.0534	5799.6187	4303.8814	4308.5192	6147.6487	6152.0794
24.5	4283.1273	4280.3012	6125.7462	6123.0445	4648.7446	4653.7229	6481.9597	6486.7132
25.5	4628.9766	4625.8634	6461.0574	6458.0832	5005.4806	5010.8024	6827.7350	6832.8133

		$\mathbf{A}^2$	Δ <sub>3/2</sub>		$\mathrm{A}^2\Delta_{5/2}$					
_	v=(	0	<b>v</b> =		v=	0	v=			
<u>J</u>	e	f	<u>e</u>	f	<u>e</u>	f	<u>e</u>	<u>f</u>		
1.5	24243.5288	24243.5288	25475.4858	25475.4858						
2.5	24249.2749	24249.2749	25480.7970	25480.7970	24267.9033	24267.9033	25498.7977	25498.7977		
3.5 4.5	24271.5245 24301.6148	24271.5245 24301.6148	25502.1638 25531.0431	25502.1638 25531.0431	24298.8250 24337.0582	24298.8250 24337.0582	25528.4502 25565.1248	25528.4502 25565.1248		
5.5	24339.3864	24339.3864	25567.2863	25567.2863	24382.7325	24382.7325	25608.9383	25608.9383		
6.5	24384.7706	24384.7706	25610.8277	25610.8277	24435.8801	24435.8801	25659.9179	25659.9179		
7.5	24437.7238	24437.7238	25661.6240	25661.6240	24496.5017	24496.5017	25718.0611	25718.0611		
8.5	24498.2101	24498.2101	25719.6386	25719.6386	24564.5834	24564.5834	25783.3513	25783.3513		
9.5	24566.1958	24566.1958	25784.8368	25784.8368	24640.1026	24640.1026	25855.7636	25855.7636		
10.5	24641.6466	24641.6466	25857.1825 25936.6376	25857.1825 25936.6376	24723.0307 24813.3342	24723.0307	25935.2671 26021.8258	25935.2671 26021.8258		
11.5 12.5	24724.5261 24814.7959	24724.5261 24814.7959	26023.1615	26023.1615	24910.9752	24813.3342 24910.9752	26115.3994	26115.3994		
13.5	24912.4145	24912.4145	26116.7106	26116.7106	25015.9120	25015.9120	26215.9433	26215.9433		
14.5	25017.3378	25017.3378	26217.2379	26217.2379	25128.0990	25128.0990	26323.4094	26323.4094		
15.5	25129.5181	25129.5181	26324.6935	26324.6935	25247.4870	25247.4870	26437.7454	26437.7454		
16.5	25248.9049	25248.9049	26439.0239	26439.0239	25374.0231	25374.0231	26558.8957	26558.8957		
17.5	25375.4444	25375.4444	26560.1722	26560.1722	25507.6510	25507.6510	26686.8007	26686.8007		
18.5	25509.0792 25649.7489	25509.0792	26688.0782 26822.6782	26688.0782 26822.6782	25648.3109 25795.9394	25648.3109 25795.9394	26821.3974 26962.6191	26821.3974 26962.6191		
19.5 20.5	25797.3894	25649.7489 25797.3894	26963.9049	26963.9049	25950.4694	25950.4694	27110.3957	27110.3957		
21.5	25951.9335	25951.9335	27111.6878	27111.6878	26111.8307	26111.8307	27264.6533	27264.6533		
22.5	26113.3102	26113.3102	27265.9525	27265.9525	26279.9493	26279.9493	27425.3146	27425.3146		
23.5	26281.4454	26281.4454	27426.6215	27426.6215	26454.7476	26454.7476	27592.2987	27592.2987		
24.5	26456.2612	26456.2612	27593.6136	27593.6136	26636.1448	26636.1448	27765.5212	27765.5212		
25.5	26637.6765	26637.6765	27766.8440	27766.8440	26824.0562	26824.0562	27944.8940	27944.8940		
26.5	26825.6064	26825.6064	27946.2247	27946.2247	27018.3936 27219.0653	27018.3936 27219.0653	28130.3258 28321.7214	28130.3258 28321.7214		
27.5 28.5	27019.9626 27220.6532	27019.9626 27220.6532	28131.6638 28323.0662	28131.6638 28323.0662	27425.9759	27425.9759	28518.9822	28518.9822		
29.5	27427.5826	27427.5826	28520.3332	28520.3332	27639.0264	27639.0264	28722.0062	28722.0062		
30.5	27640.6518	27640.6518	28723.3625	28723.3625	27858.1142	27858.1142	28930.6878	28930.6878		
_	·····	X <sup>2</sup> I	T	<u></u>		X <sup>2</sup> ]		· · · · · · · · · · · · · · · · · · ·		
0.5	62 6772	-63.6290	1368.9821	1369.0307			3/2			
0.5 1.5	-63.6773 -52.4937	-52.3983	1379.9377	1389.0307	83.2427	83.2437	1516.2374	1516.2383		
2.5	-33.8380	-33.6975	1398.2135	1398.3552	103.0143	103.0182	1535.5806	1535.5844		
3.5	-7.7104	-7.5277	1423.8088	1423.9930	130.6762	130.6859	1562.6439	1562.6531		
4.5	25.8882	26.1093	1456.7224	1456.9453	166.2111	166.2302	1597.4107	1597.4287		
5.5	66.9558	67.2109	1496.9519	1497.2088	209.5973	209.6301	1639.8606	1639.8915		
6.5	115.4893	115.7732	1544.4938	1544.7795	260.8097	260.8609	1689.9694	1690.0179		
7.5 8.5	171.4834	171.7906 235.2558	1599.3425 1661.4909	1599.6512 1661.8165	319.8193 386.5944	319.8942 386.6985	1747.7100 1813.0520	1747.7810 1813.1507		
8.3 9.5	234.9310 305.8226	306.1589	1730.9295	1731.2655	461.1006	461.2394	1885.9627	1886.0946		
10.5	384.1460	384.4877	1807.6463	1807.9861	543.3007	543.4800	1966.4070	1966.5776		
11.5	469.8862	470.2272	1891.6272	1891.9638	633.1556	633.3812	2054.3476	2054.5627		
12.5	563.0254	563.3593	1982.8551	1983.1817	730.6241	730.9016	2149.7458	2150.0108		
13.5	663.5426	663.8635	2081.3103	2081.6199	835.6630	835.9981	2252.5607	2252.8813		
14.5	771.4138	771.7156	2186.9707	2187.2563	948.2274	948.6254	2362.7504	2363.1319		
15.5	886.6120	886.8890	2299.8112	2300.0660	1068.2706 1195.7440	1068.7366	2480.2713	2480.7190		
16.5 17.5	1009.1072 1138.8662	1009.3535 1139.0763	2419.8045 2546.9205	2420.0214 2547.0926	1330.5977	1196.2832 1331.2148	2605.0784 2737.1258	2605.5975 2737.7213		
18.5	1275.8530	1276.0216	2681.1271	2681.2476	1472.7797	1473.4794	2876.3662	2877.0430		
19.5	1420.0287	1420.1506	2822.3896	2822.4515	1622.2366	1623.0234	3022.7513	3023.5142		
20.5	1571.3516	1571.4218	2970.6713	2970.6680	1778.9133	1779.7912	3176.2320	3177.0854		
21.5	1729.7774	1729.7910	3125.9337	3125.8582	1942.7530	1943.7260	3336.7579	3337.7064		
22.5	1895.2588	1895.2113	3288.1361	3287.9818	2113.6973	2114.7691	3504.2780	3505.3259		
23.5	2067.7464	2067.6333	3457.2364	3456.9965	2291.6860	2292.8601	3678.7407	3679.8922		
24.5	2247.1880	2247.0049 2433.2721	3633.1909 3815.9546	3632.8588 3815.5234	2476.6574 2668.5481	2477.9371 2669.9365	3860.0935 4048.2833	3861.3525 4049.6539		
25.5 26.5	2433.5291 2626.7130	2433.2721 2626.3781	4005.4811	3813.3234 4004.9442	2867.2933	2868.7931	4243.2568	4049.0339		
27.5	2826.6808	2826.2642	4201.7231	4201.0738	3072.8264	3074.4403	4444.9603	4446.5653		
28.5	3033.3712	3032.8694	4404.6325	4403.8640	3285.0793	3286.8097	4653.3397	4655.0674		
29.5	3246.7211	3246.1307	4614.1603	4613.2659	3503.9824	3505.8314	4868.3410	4870.1949		
30.5	3466.6654	3465.9832	4830.2572	4829.2302	3729.4648	3731.4344	5089.9102	5091.8938		

predissociation. We are unable to add much to these observations since our 1–1 band is weak in intensity and the 2–2 band is almost absent in our spectra.

# CONCLUSION

The emission spectra of the  $A^2\Delta - X^2\Pi$  electronic transition of SiH and SiD have been measured with improved precision. The high-resolution measurements from spectra recorded with a Fourier transform spectrometer have been combined with the previous infrared vibration–rotation measurements (35–38) to extract improved molecular constants for the ground and excited states of SiH and SiD.

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## REFERENCES

- C. Jascheck and M. Jascheck, "The Behavior of Chemical Elements in Stars," Cambridge Press, Cambridge, UK, 1995.
- J. M. Jasinski, R. Becerra, and R. Walsh, Chem. Rev. 95, 1203–1228 (1995).
- G. Turban, Y. Catherine, and B. Grolleau, *Thin Solid Films* 67, 309–320 (1980), 77, 287–300 (1981).
- B. Drevillon, J. Huc, A. Lloret, J. Perrin, G. de Rosny, and J. P. M. Schmitt, Appl. Phys. Lett. 37, 646–648 (1980).
- 5. A. E. Douglas and G. A. Elliott, Can. J. Phys. 43, 496-502 (1965).
- 6. A. Schadee, Bull. Astron. Inst. Neth. 17, 311-357 (1964).
- 7. A. J. Sauval, Solar Phys. 10, 319-329 (1969).
- D. L. Lambert and E. A. Mallia, Mon. Not. R. Astr. Soc. 148, 313–324 (1970).
- M. Taniguchi, M. Hirose, T. Hamasaki, and Y. Osaka, Appl. Phys. Lett. 37, 787–788 (1980).
- A. Matsuda, K. Nakagawa, K. Tanaka, M. Matsumura, S. Yamasaki, H. Okushi, and S. Iizima, J. Non-Cryst. Solids 35/36, 183–188 (1980).
- 11. J. Perrin and E. Delafosse, J. Phys. D 13, 759-765 (1980).
- 12. F. J. Kampas and R. E. Griffith, J. Appl. Phys. 52, 1285-1288 (1981).
- 13. J. Perrin and J. P. M. Schmitt, Chem. Phys. 67, 167-176 (1982).
- J. P. M. Schmitt, P. Gressier, M. Krishnan, G. de Rosny, and J. Perrin, Chem. Phys. 84, 281–293 (1984).
- Y. Matsumi, T. Hayashi, H. Yoshikawa, and S. Komiya, J. Vac. Sci. Technol. A 4, 1786–1790 (1986).
- 16. W. G. Tong and R. W. Shaw, Appl. Spectrosc. 40, 494-497 (1986).
- 17. C. V. Jackson, Proc. Roy. Soc. A 126, 373-392 (1930).
- 18. R. S. Mulliken, Phys. Rev. 37, 733-735 (1931).
- 19. G. D. Rochester, Z. Physik 101, 769-784 (1936).
- 20. A. E. Douglas, Can. J. Phys. 35, 71-77 (1957).
- 21. L. Klynning and B. Lindgren, Ark. Fysik 33, 73-91 (1966).
- 22. R. D. Verma, Can. J. Phys. 43, 2136-2141 (1965).

- G. Herzberg, A. Lagerqvist, and B. J. McKenzie, Can. J. Phys. 47, 1889–1897 (1969).
- 24. P. Bollmark, L. Klynning, and P. Pagès, Phys. Scr. 3, 219-222 (1971).
- 25. R. D. Johnson III and J. W. Hudgens, J. Phys. Chem. 93, 6268–6270 (1989).
- S. Weinreb, A. H. Barrett, M. L. Meeks, and J. C. Henry, *Nature* 200, 829–831 (1963).
- O. E. H. Rydbeck, I. Elldér, and W. M. Irwine, *Nature* 246, 466–486 (1973).
- 28. I. D. L. Wilson and W. G. Richards, Nature 258, 133-134 (1975).
- 29. L. Klynning, B. Lindgren, and U. Sassenberg, Phys. Scr. 20, 617–619 (1979).
- 30. D. L. Cooper and W. G. Richards, J. Chem. Phys. 74, 96-98 (1981).
- 31. R. S. Freedman and A. W. Irwin, Astron. Astrophys. 53, 447–449 (1976).
- R. N. Zare, A. L. Schmeltekopf, W. J. Harrop, and D. L. Albritton, *J. Mol. Spectrosc.* 46, 37–66 (1973).
- J. C. Knights, J. P. M. Schmitt, J. Perrin, and G. Guelachvili, J. Chem. Phys. 76, 3414–3421 (1982).
- 34. J. M. Brown and D. Robinson, Mol. Phys. 51, 883-886 (1984).
- J. M. Brown, R. F. Curl, and K. M. Evenson, J. Chem. Phys. 81, 2884–2890 (1984).
- P. B. Davies, N. A. Isaacs, S. A. Johnson, and D. K. Russell, *J. Chem. Phys.* 83, 2060–2063 (1985).
- M. Betrencourt, D. Boudjaadar, P. Chollet, G. Guelachvili, and M. Morillon-Chapey, J. Chem. Phys. 84, 4121–4126 (1986).
- W. Seebass, J. Werner, W. Urban, E. R. Comben, and J. M. Brown, Mol. Phys. 62, 161–174 (1987).
- 39. C. Park, J. Quant. Spectrosc. Radiat. Transfer 21, 373–385 (1979).
- N. Washida, Y. Matsumi, T. Hayashi, T. Ibuki, A. Hiraya, and K. Shobatake, J. Chem. Phys. 83, 2769–2774 (1985).
- 41. S. Stamou, D. Mataras, and D. Rapakoulias, Chem. Phys. 218, 57–69 (1997).
- 42. N. Grevesse and A. J. Sauval, J. Quant. Spectrosc. Radiat. Transfer 11, 65–67 (1971).
- 43. W. H. Smith, J. Chem. Phys. 51, 520-524 (1969).
- 44. W. H. Smith and H. S. Liszt, J. Quant. Spectrosc. Radiat. Transfer 11, 45–54 (1971).
- T. A. Carlson, N. Duri, P. Erman, and M. Larsson, J. Phys. B 11, 3667–3675 (1978).
- W. Bauer, K. H. Becker, R. Düren, C. Hubrich, and R. Meuser, *Chem. Phys. Lett.* 108, 560–561 (1984).
- L. G. M. Pettersson and S. R. Langhoff, *Chem. Phys. Lett.* 125, 429–432 (1986).
- M. Lewerenz, P. J. Bruna, S. D. Peyerimhoff, and R. J. Buenker, *Mol. Phys.* 49, 1–24 (1983).
- 49. W. Meyer and P. Rosmus, J. Chem. Phys. 63, 2356-2375 (1975).
- 50. J. K. Park and H. Sun, Chem. Phys. Lett. 195, 469-474 (1992).
- 51. A. F. Sax and J. Kalcher, J. Phys. Chem. 95, 1768-1783 (1991).
- 52. L. A. Curtiss and J. A. Pople, Chem. Phys. Lett. 144, 38-42 (1988).
- 53. M. Larsson, J. Chem. Phys. 86, 5018-5026 (1987).
- P. Ho, M. E. Coltrin, J. S. Binkley, and C. F. Melius, J. Phys. Chem. 89, 4647–4654 (1985).
- 55. J. Kalcher, Chem. Phys. 118, 273-284 (1987).
- R. A. Keller, B. E. Warner, E. F. Zaleweski, P. Dyer, R. Engleman, Jr., and
   B. A. Palmer, J. Physique Collog. 44, C7–C23 (1983).
- B. A. Palmer and R. Engleman, Jr., "Atlas of Thorium Spectrum." Report LA-9615, Los Alamos National Laboratory, Los Alamos, NM, 1983.
- J. M. Brown, E. A. Colbourn, J. K. G. Watson, and F. D. Wayne, J. Mol. Spectrosc. 74, 294–318 (1979).
- C. Amiot, J.-P. Maillard, and J. Chauville, J. Mol. Spectrosc. 87, 196–218
- J. M. Brown, A. S-C. Cheung, and A. J. Merer, J. Mol. Spectrosc. 124, 464–475 (1987).
- 61. J. M. Brown, private communication.
- 62. J. M. Brown and J. K. G. Watson, J. Mol. Spectrosc. 65, 65-74 (1977).